

A brief history of the universe – and how we contribute

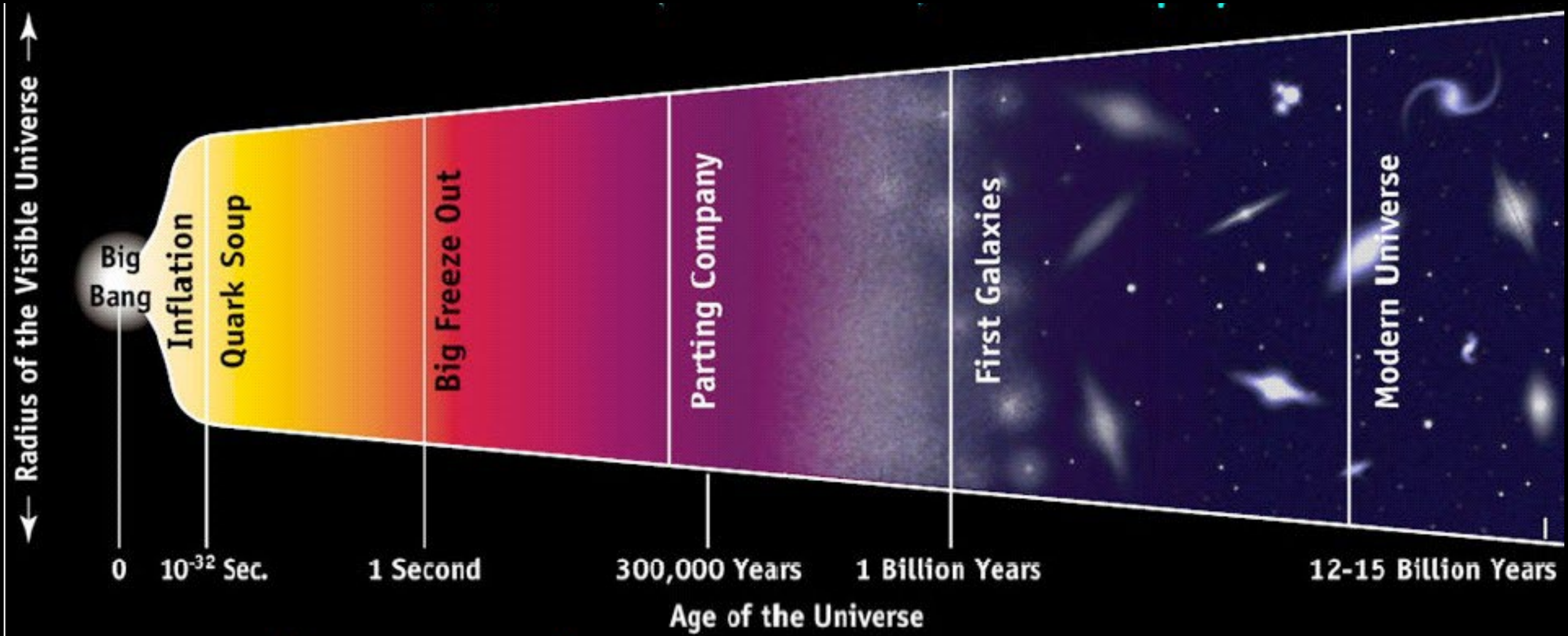
René Reifarth

P-3, LANL

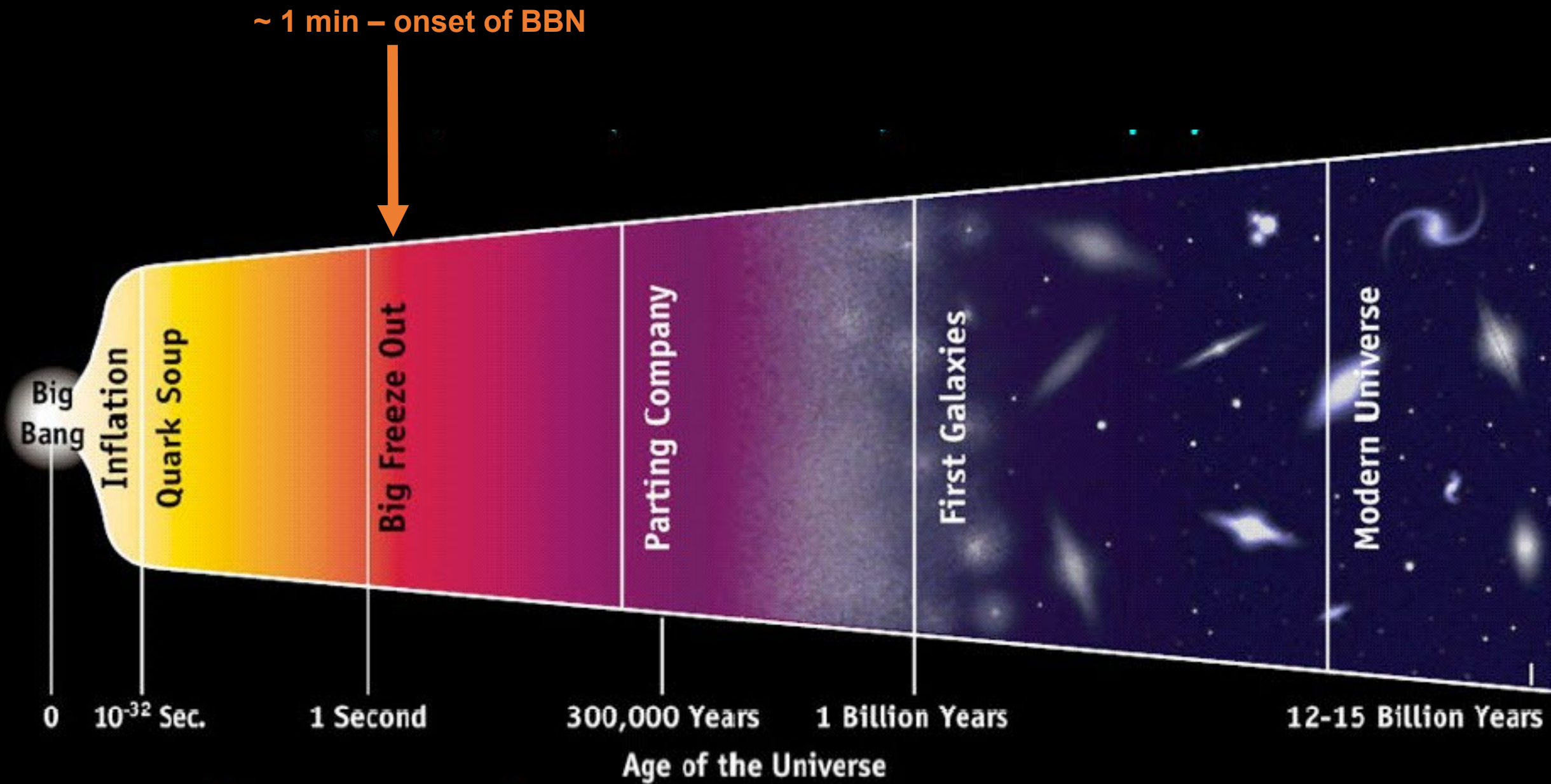
P-3 seminar

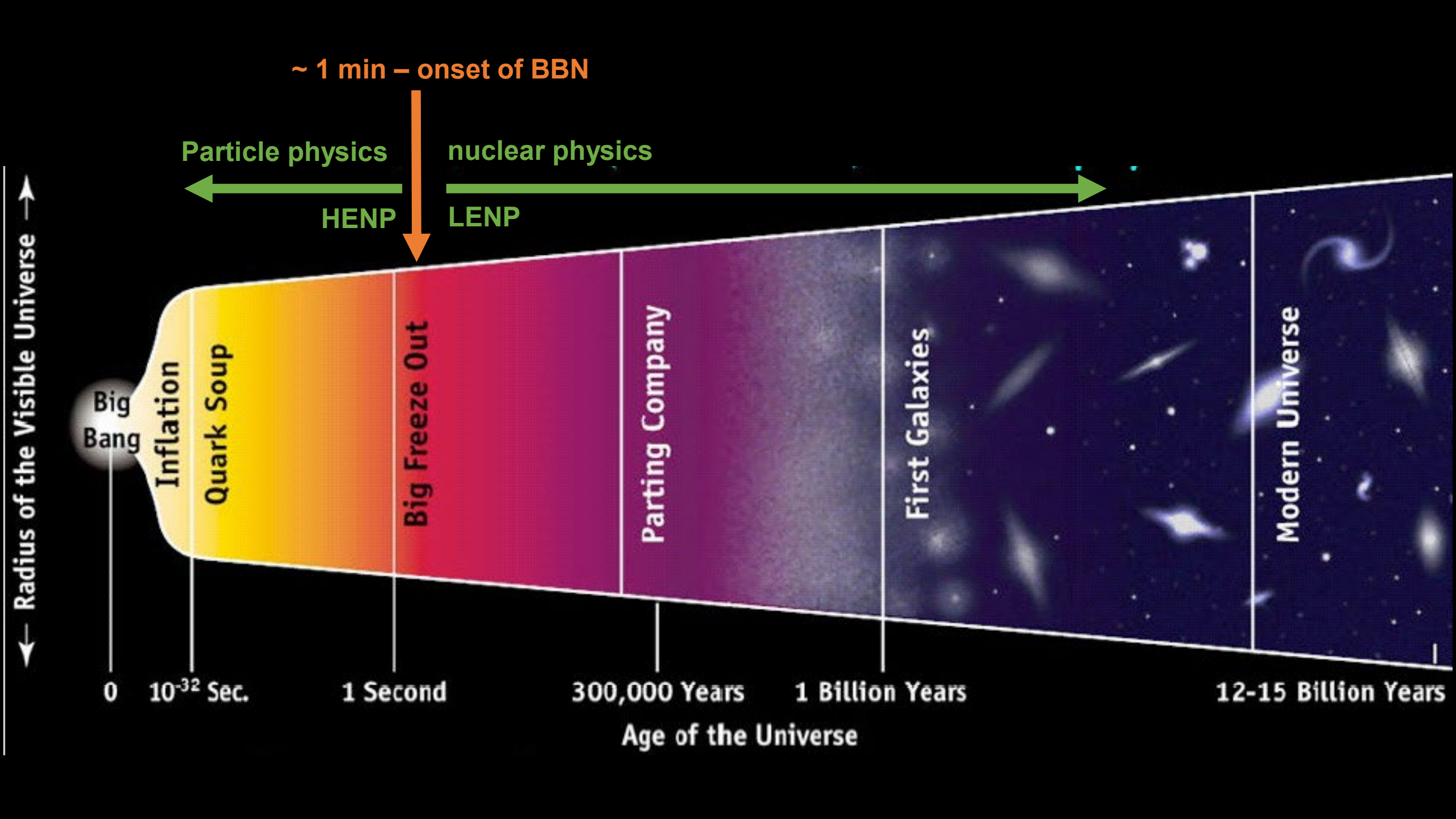
1. November 2023, LANSCE, LANL, USA

The history of the universe – as we believe it

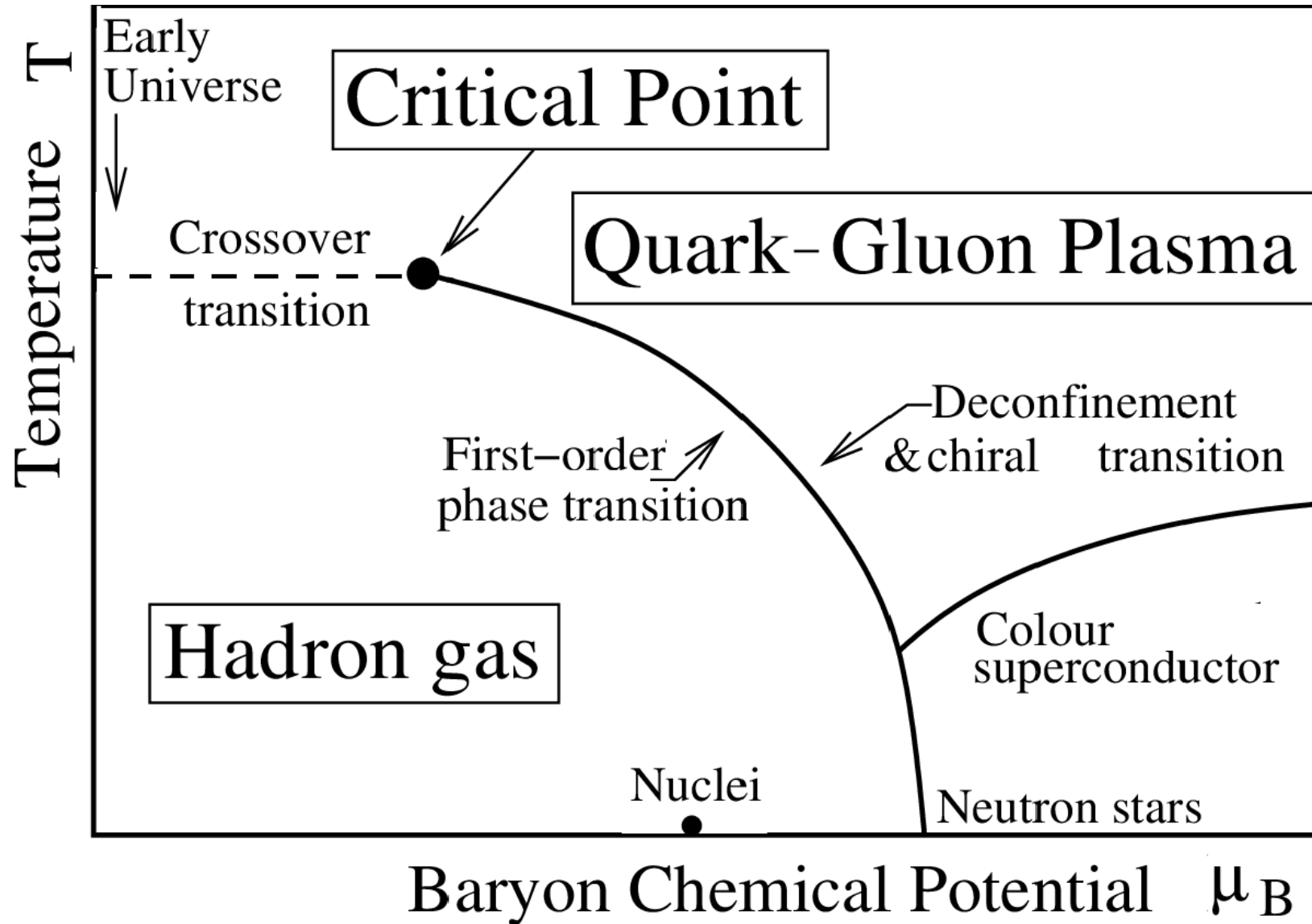


← Radius of the Visible Universe →

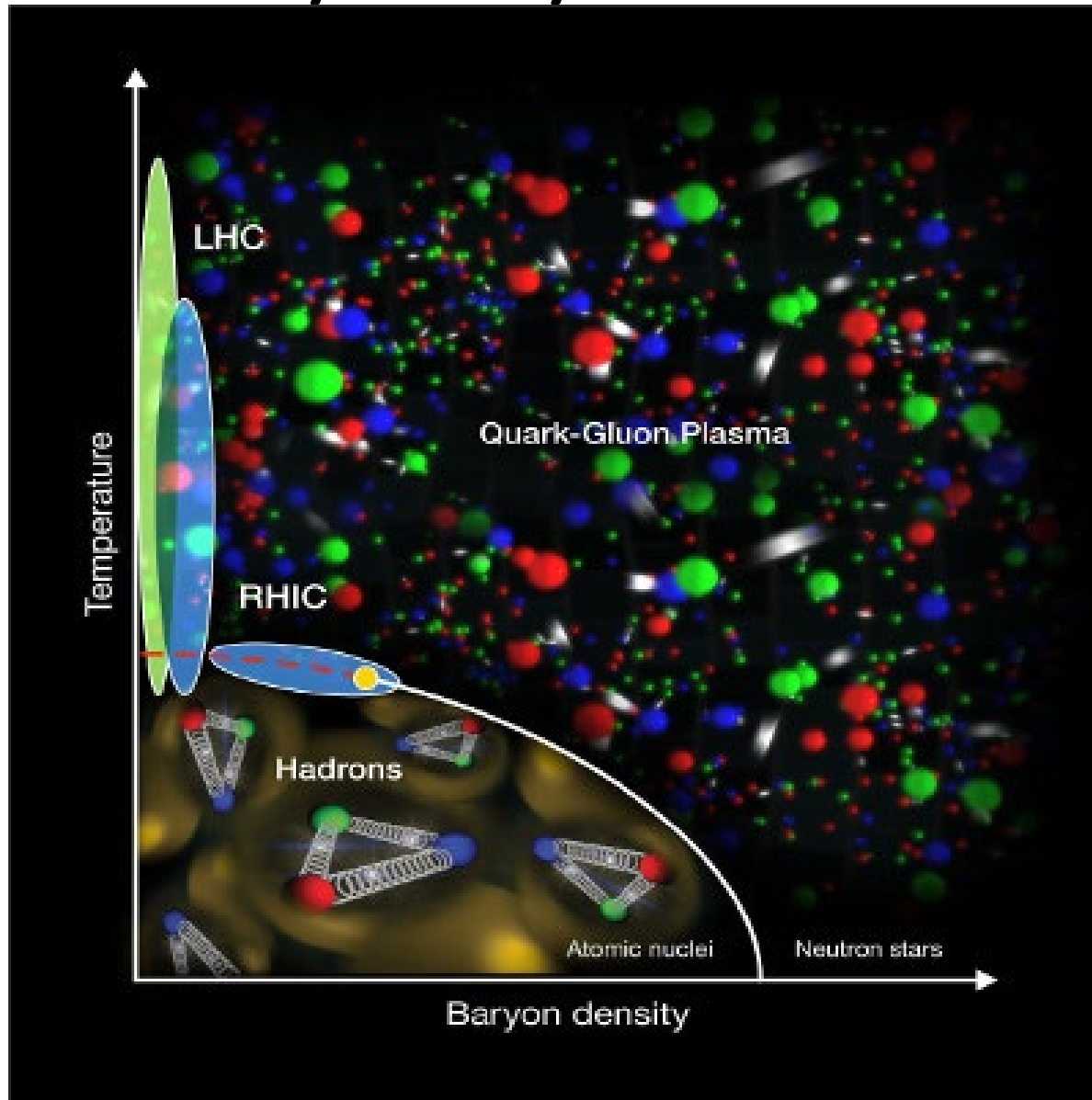




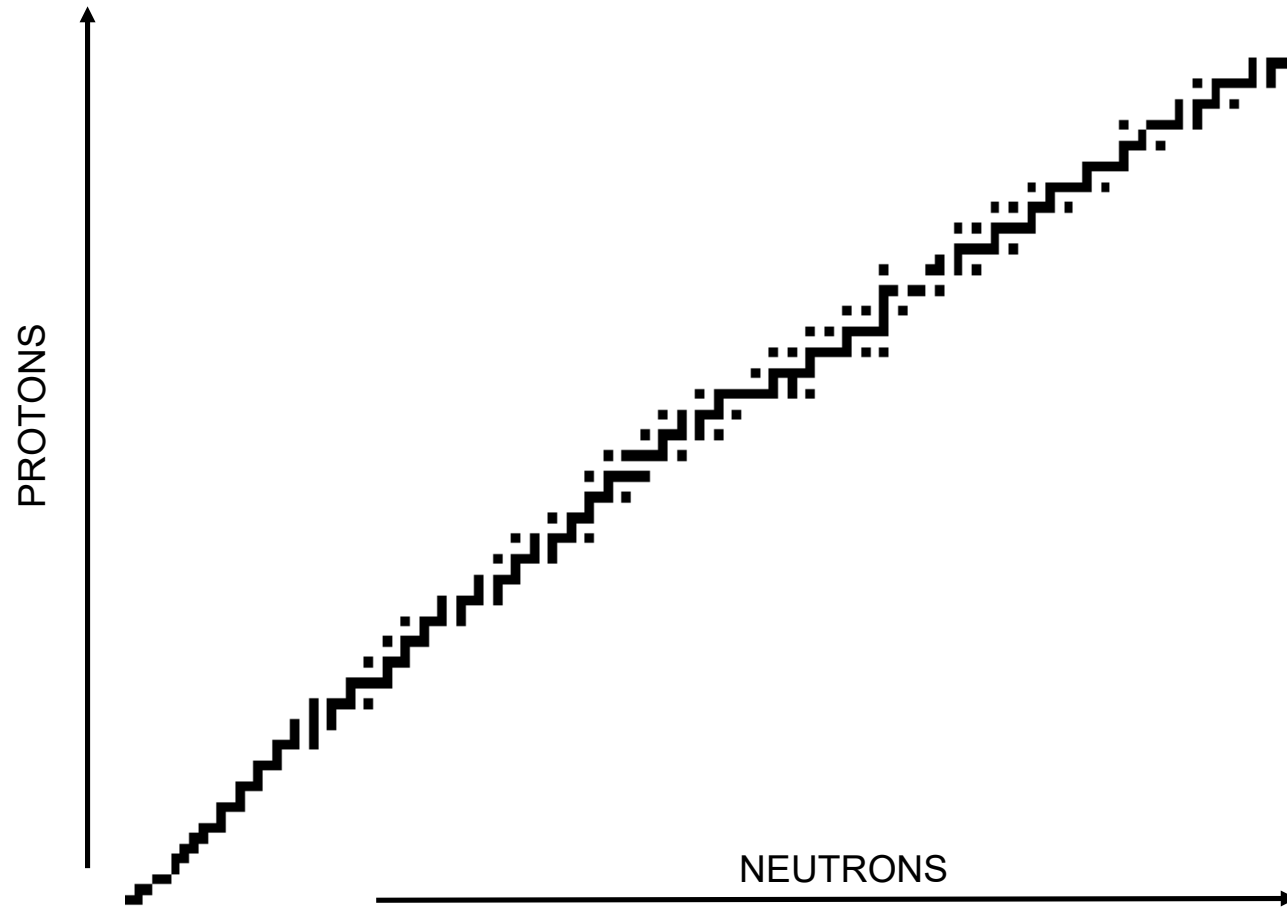
The very early universe



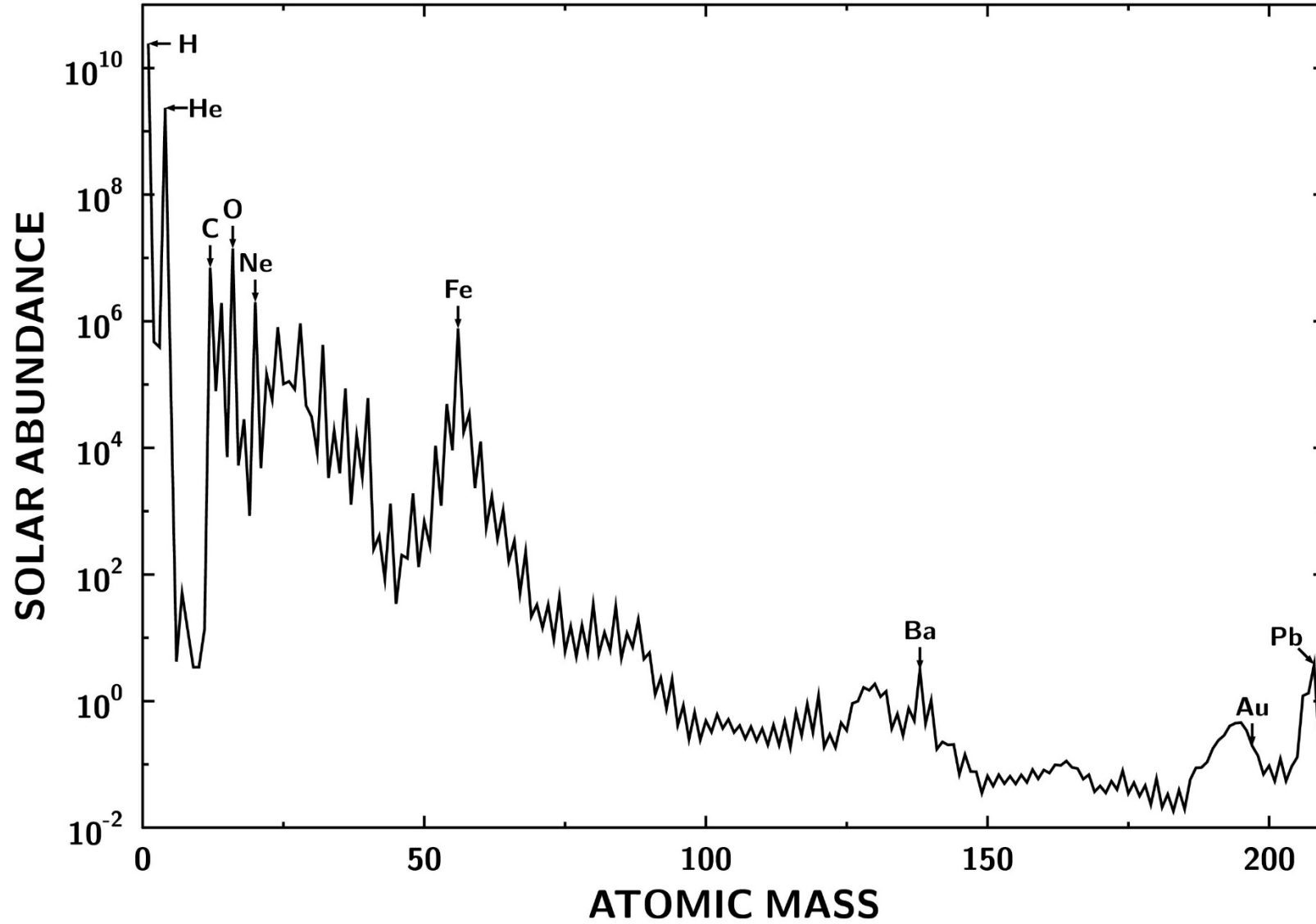
The very early universe - HENP



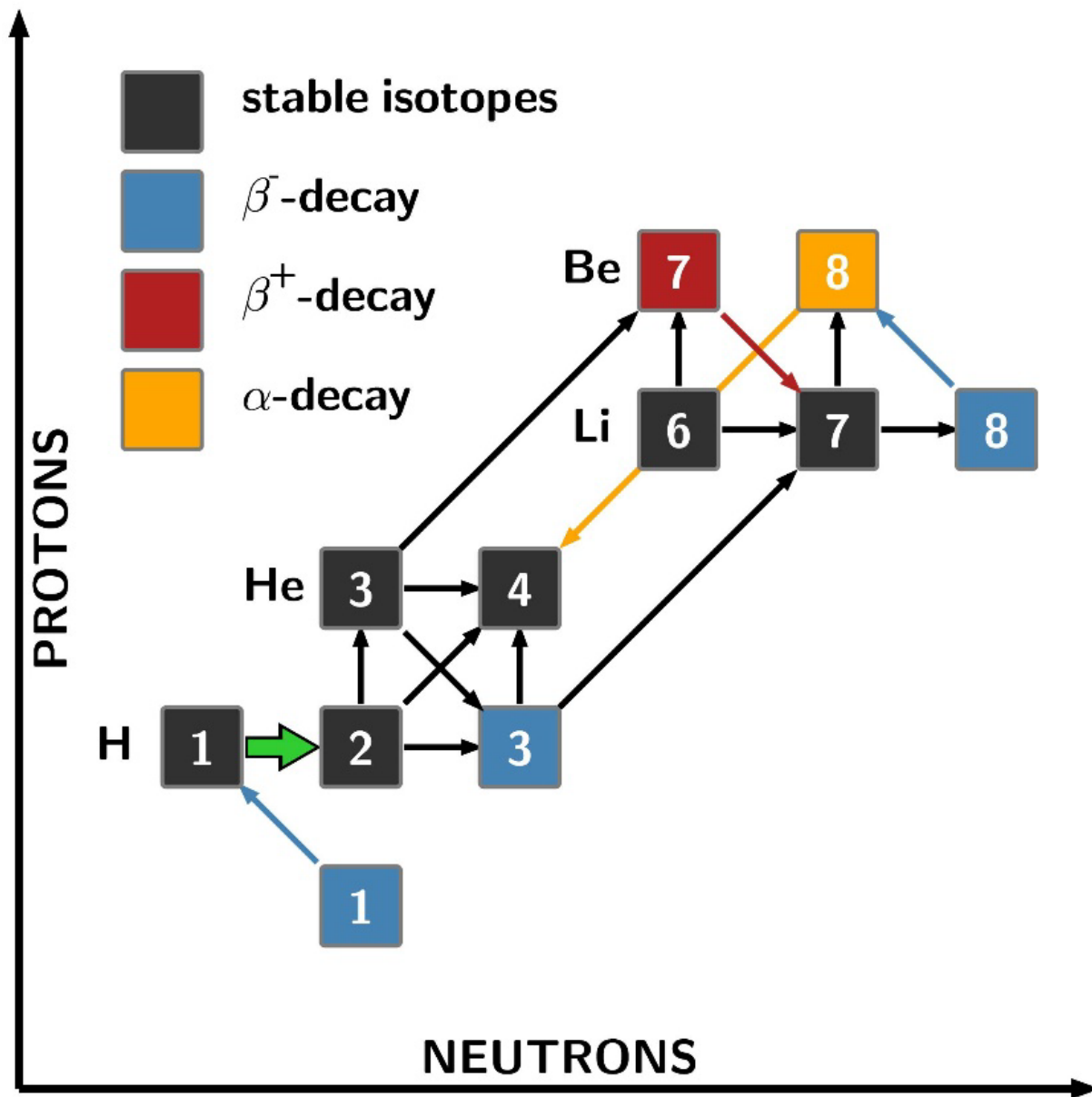
The chart of stable nuclei



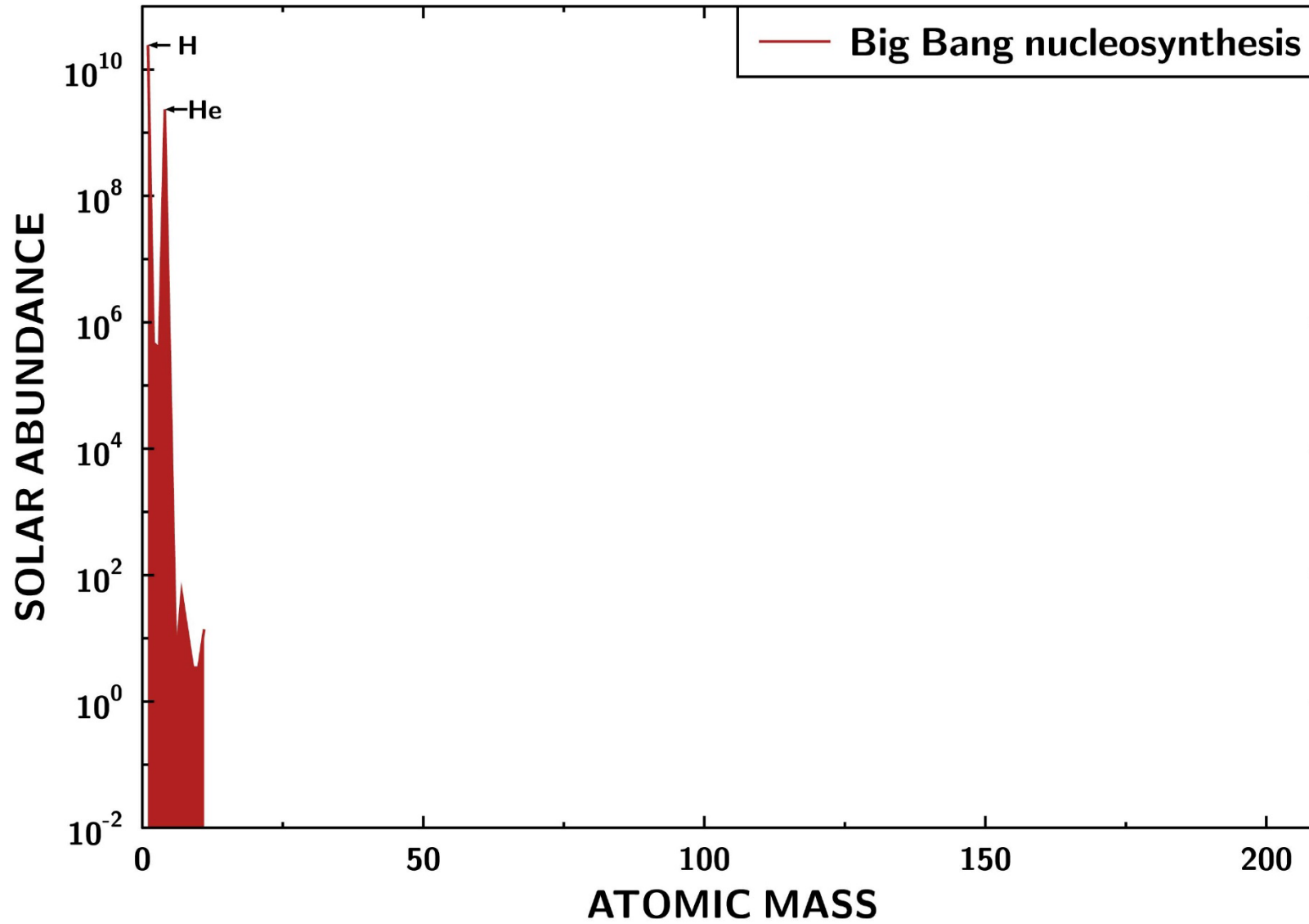
Solar abundances



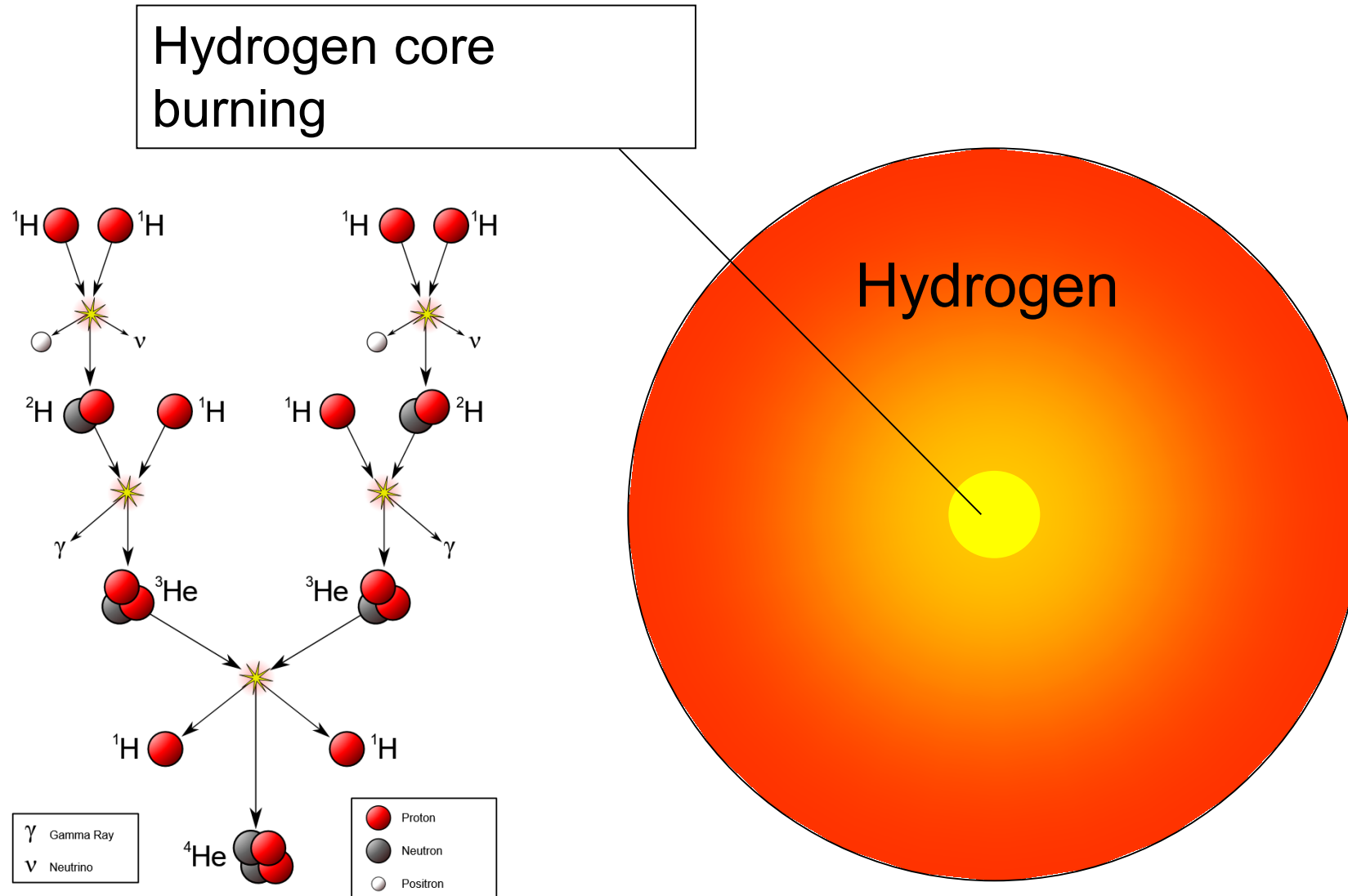
Big Bang Nucleosynthesis



First elements – H, He

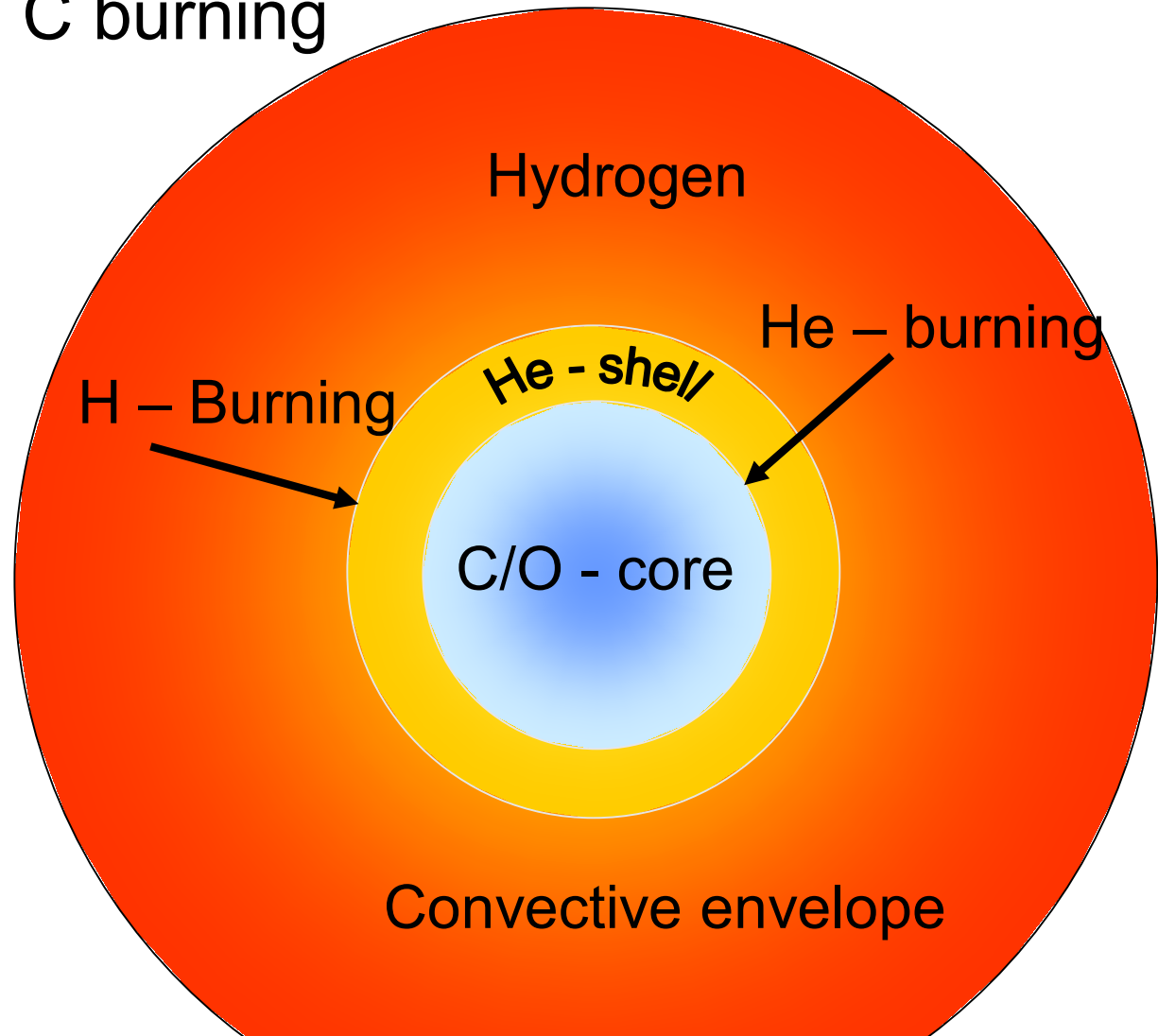


First stars after 500 million years

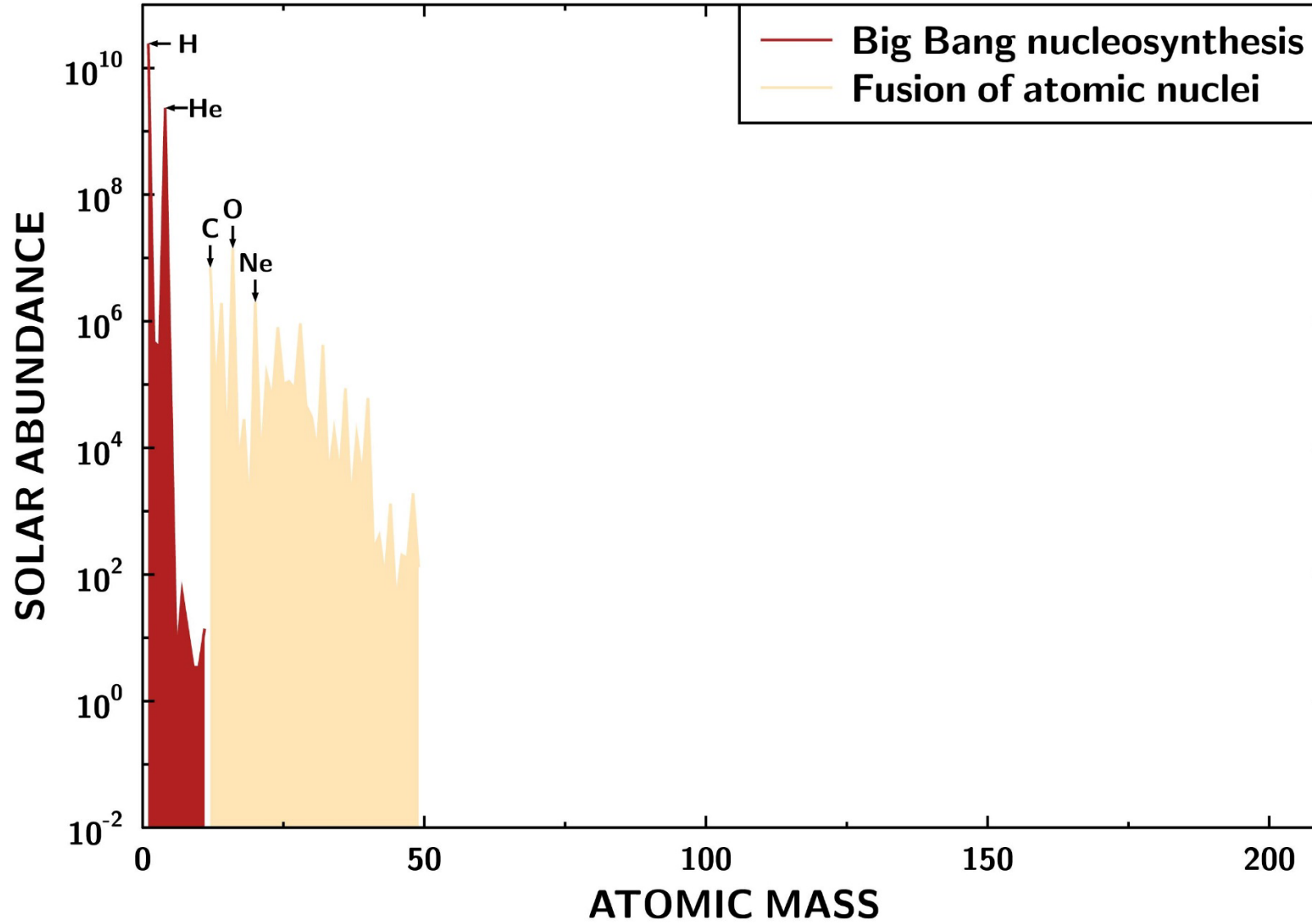


Onion structure

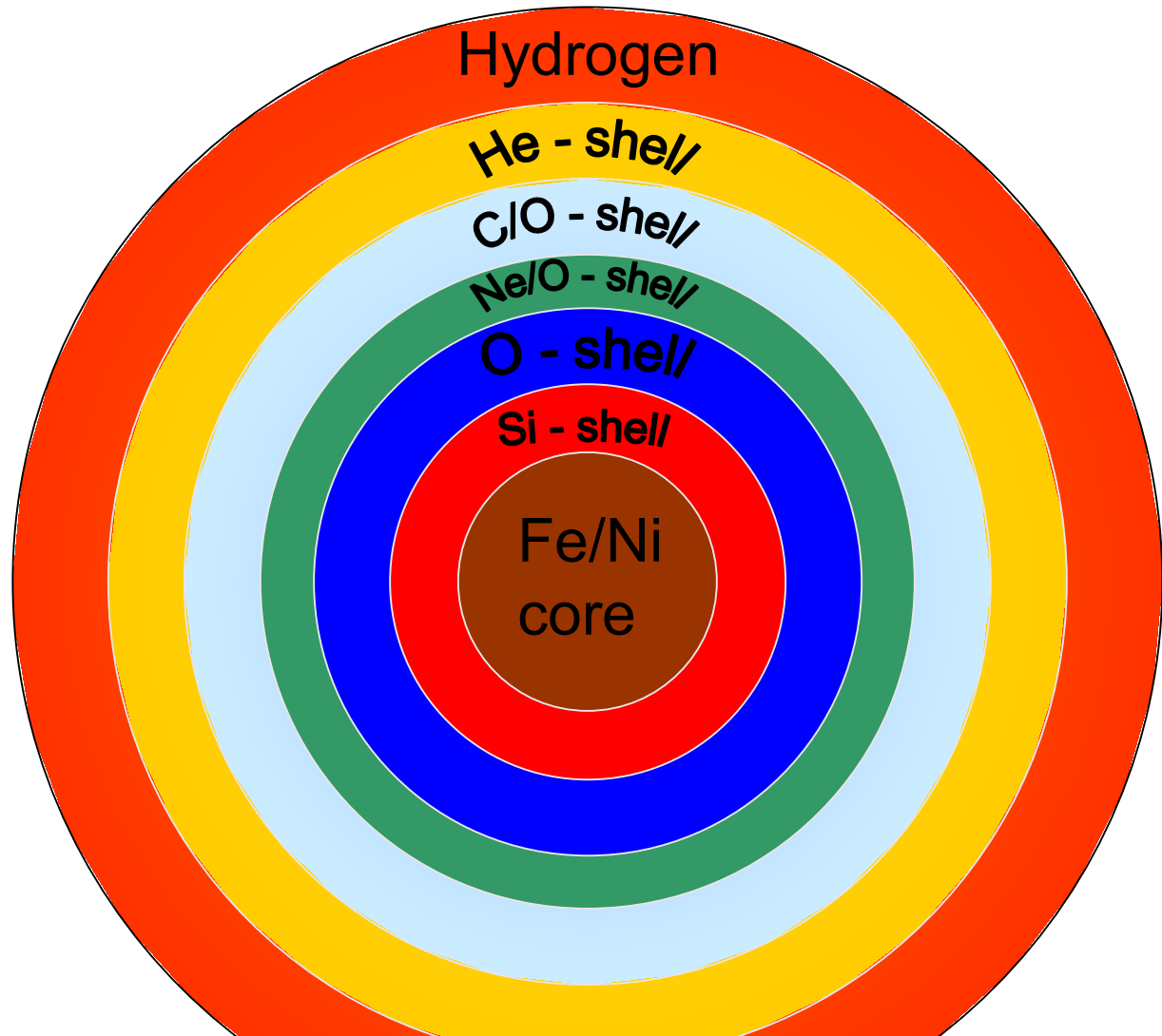
Later stages: H, He, C burning



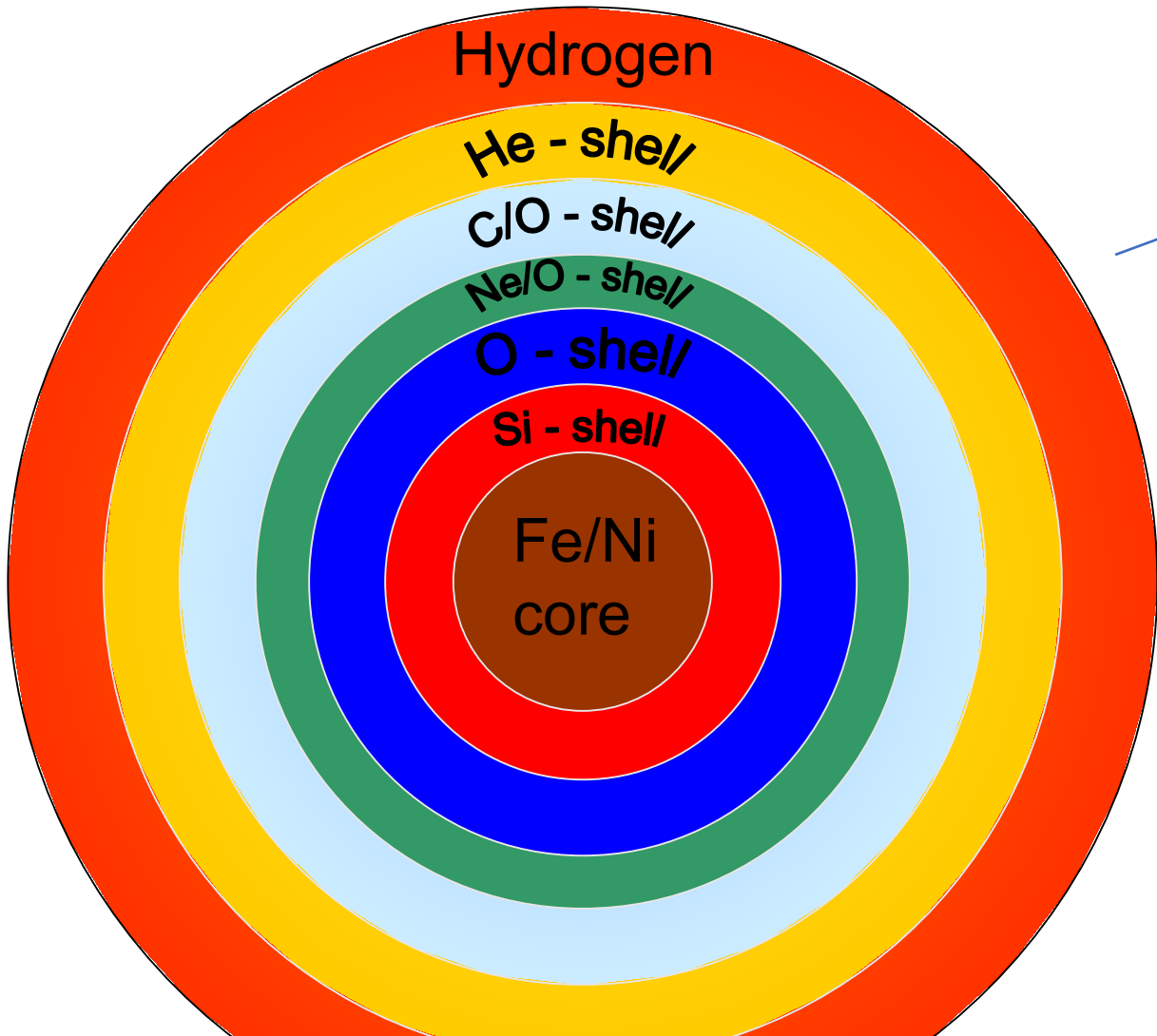
Energy source of stars



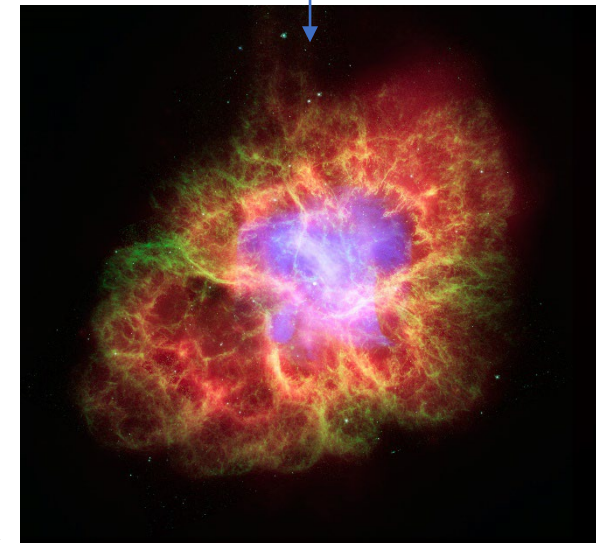
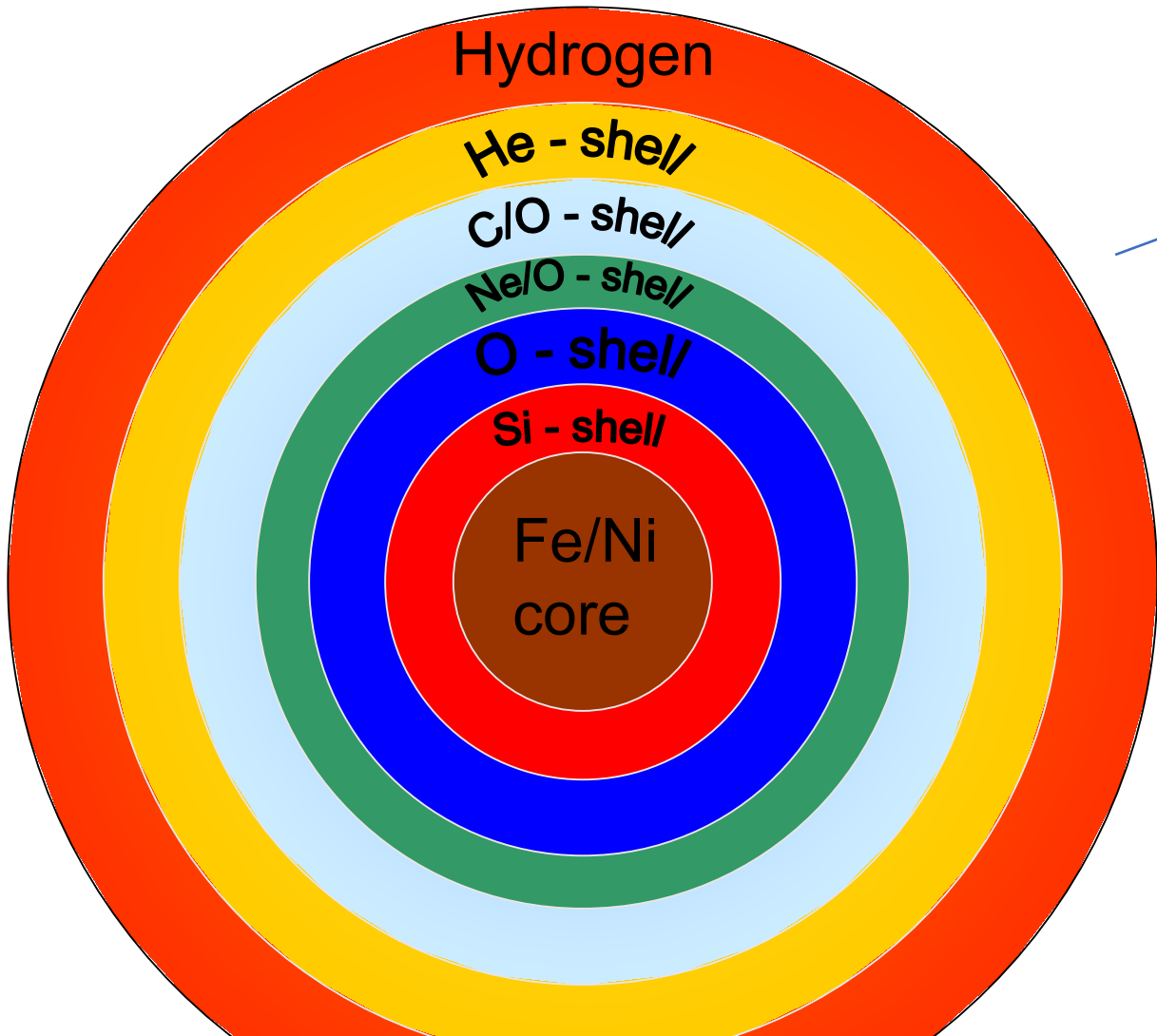
Massive stars – early death



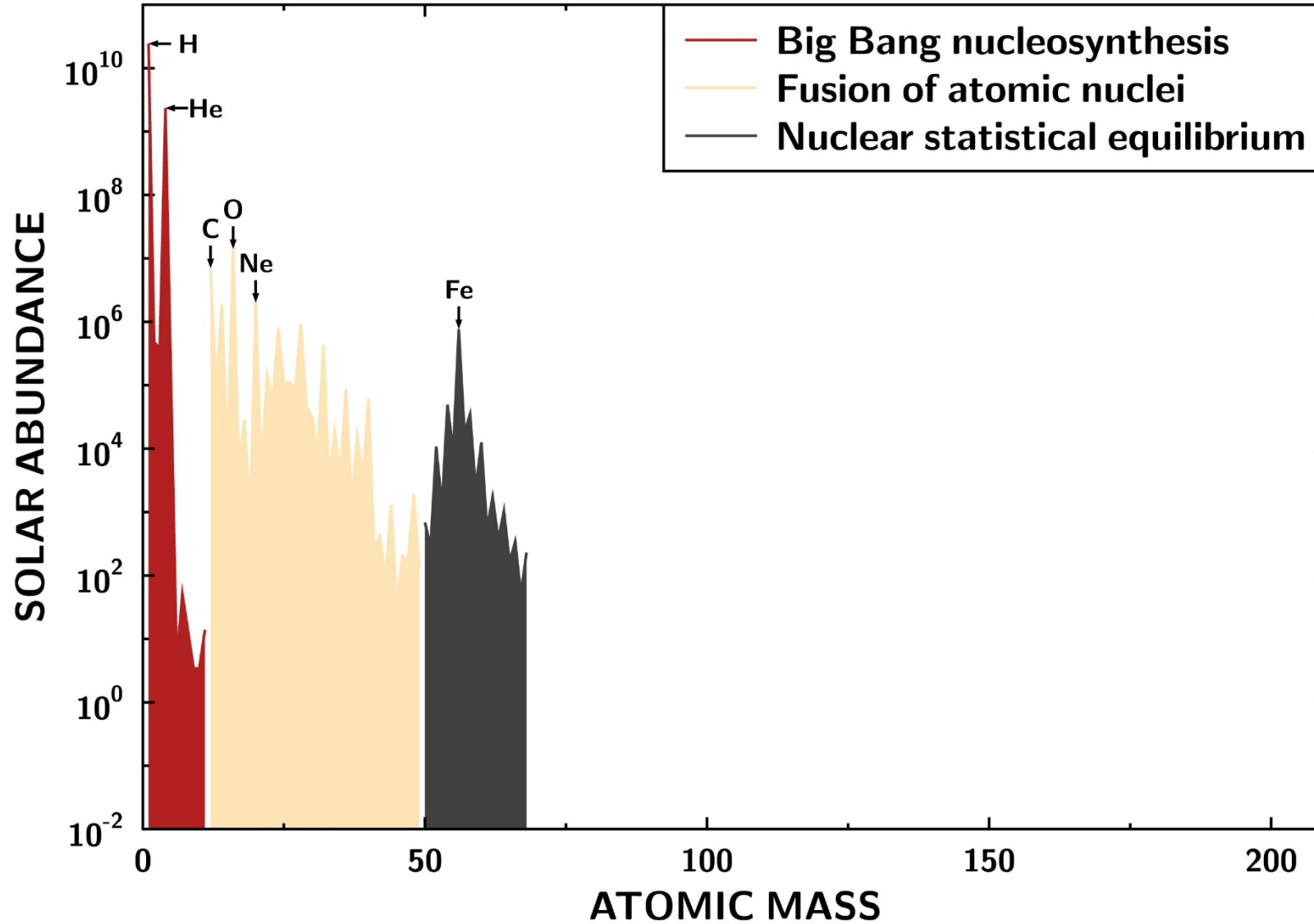
Massive stars – early death



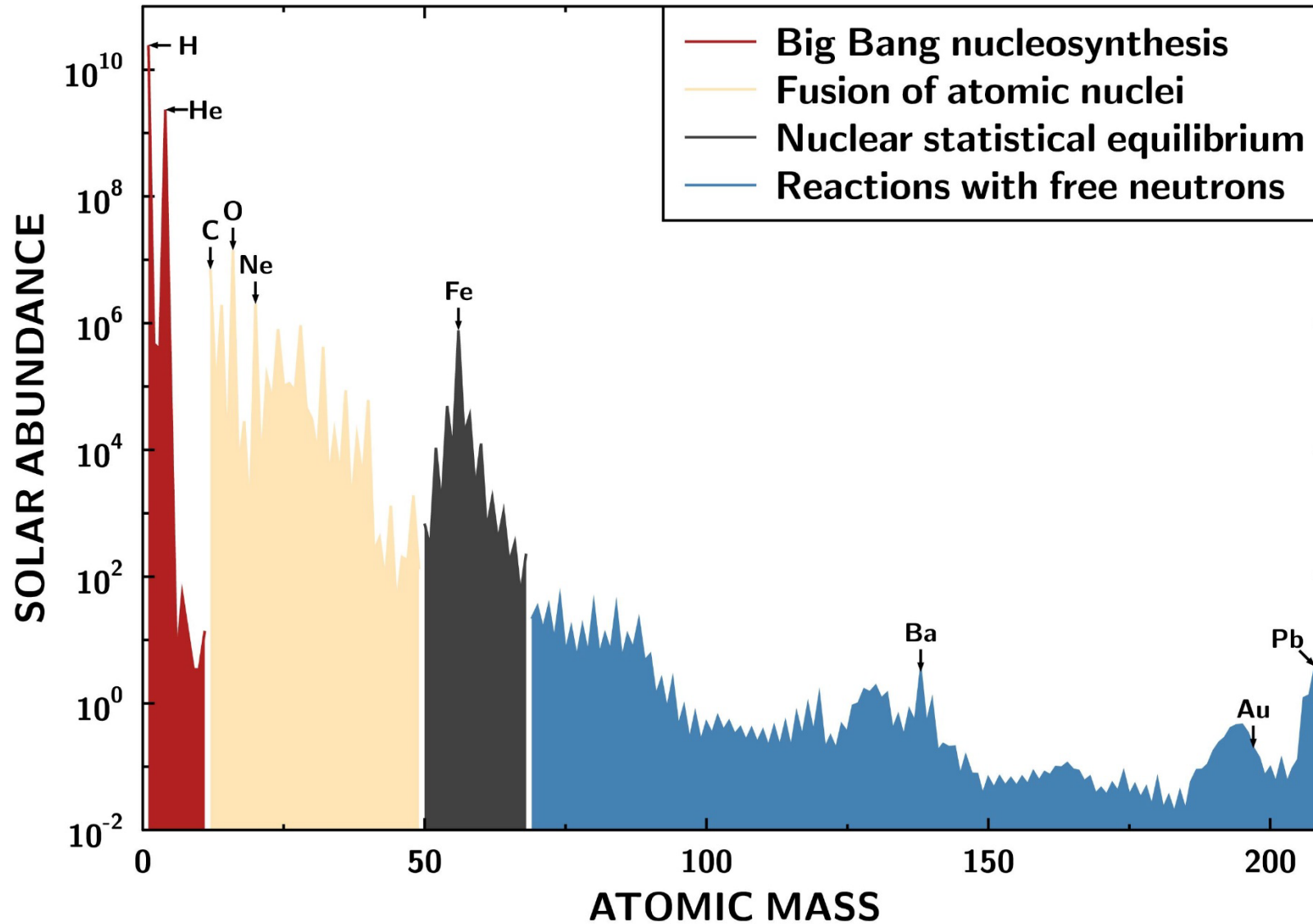
Massive stars – early death



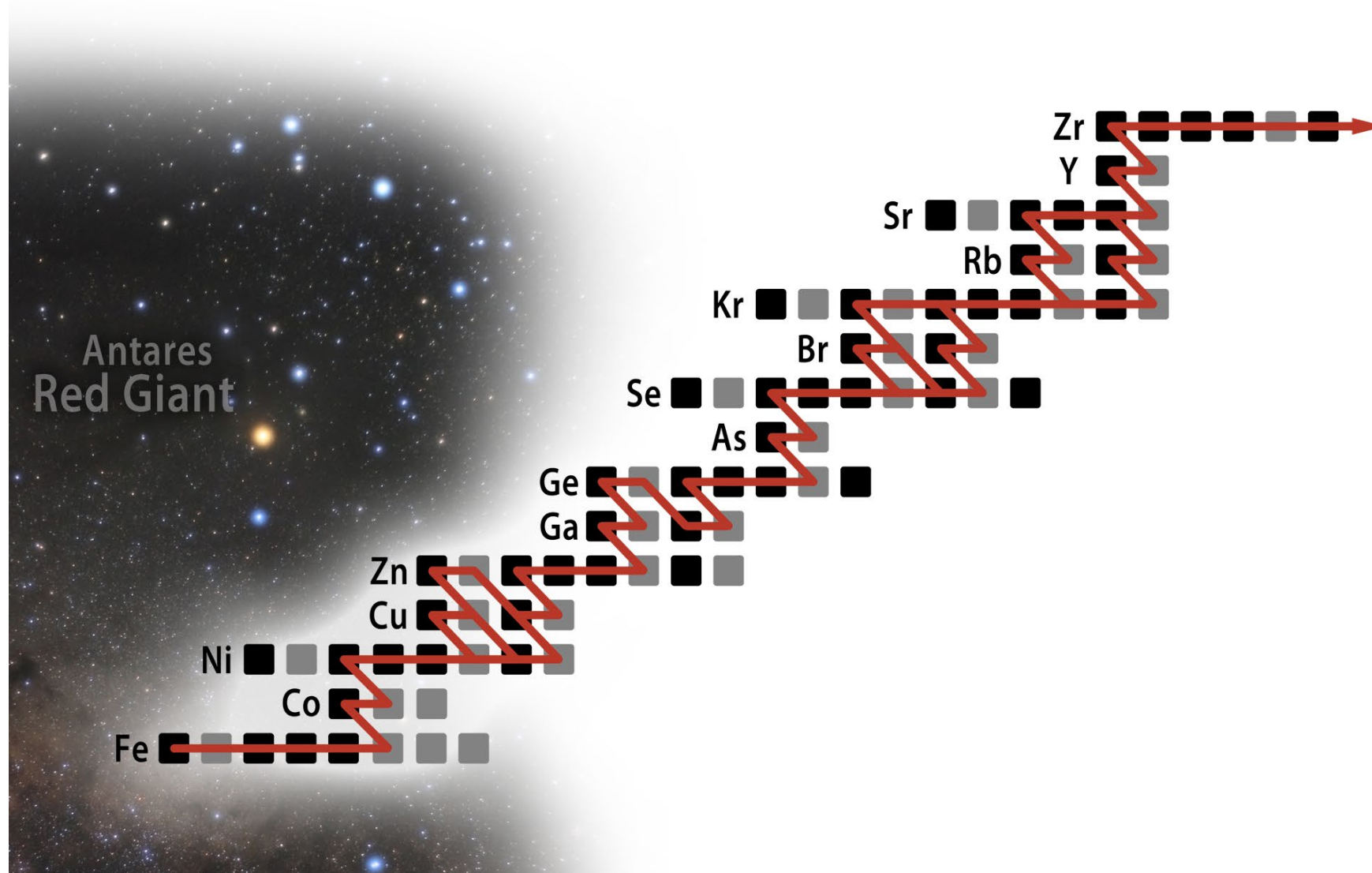
Iron – survival of the most stable



The synthesis of the elements



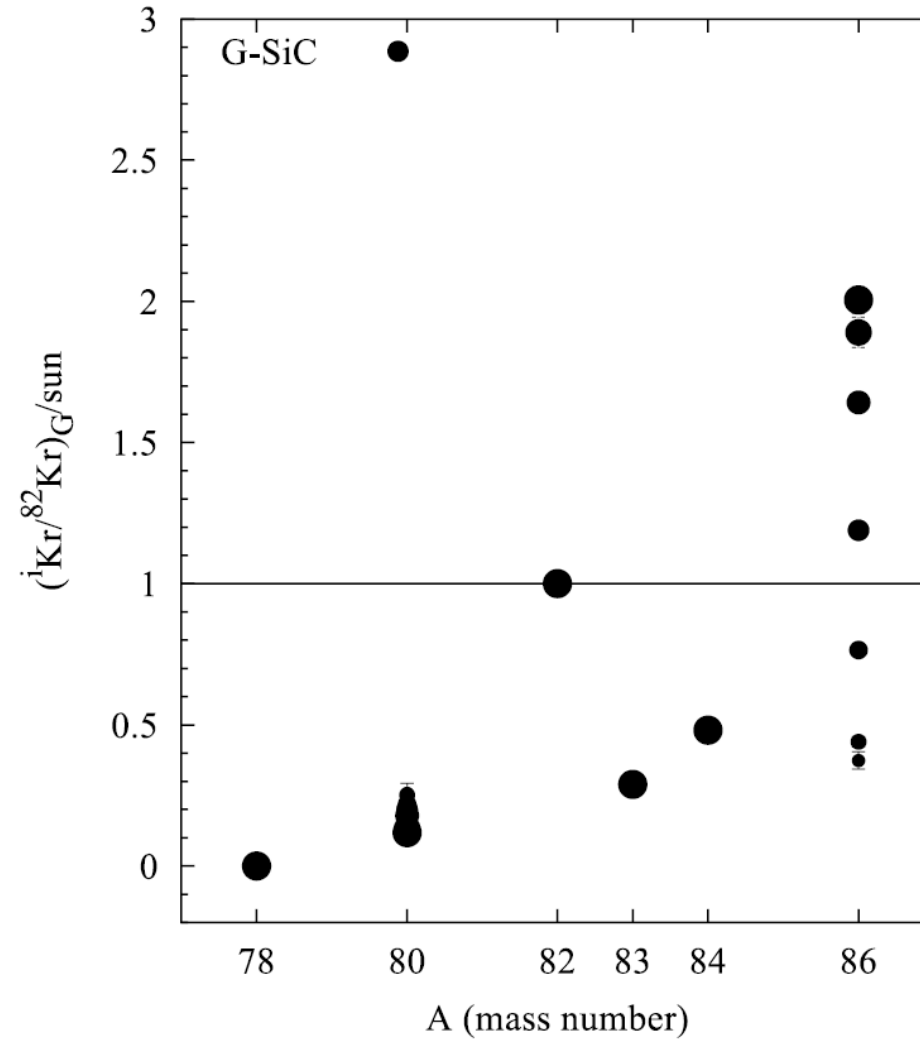
Neutron-induced nucleosynthesis



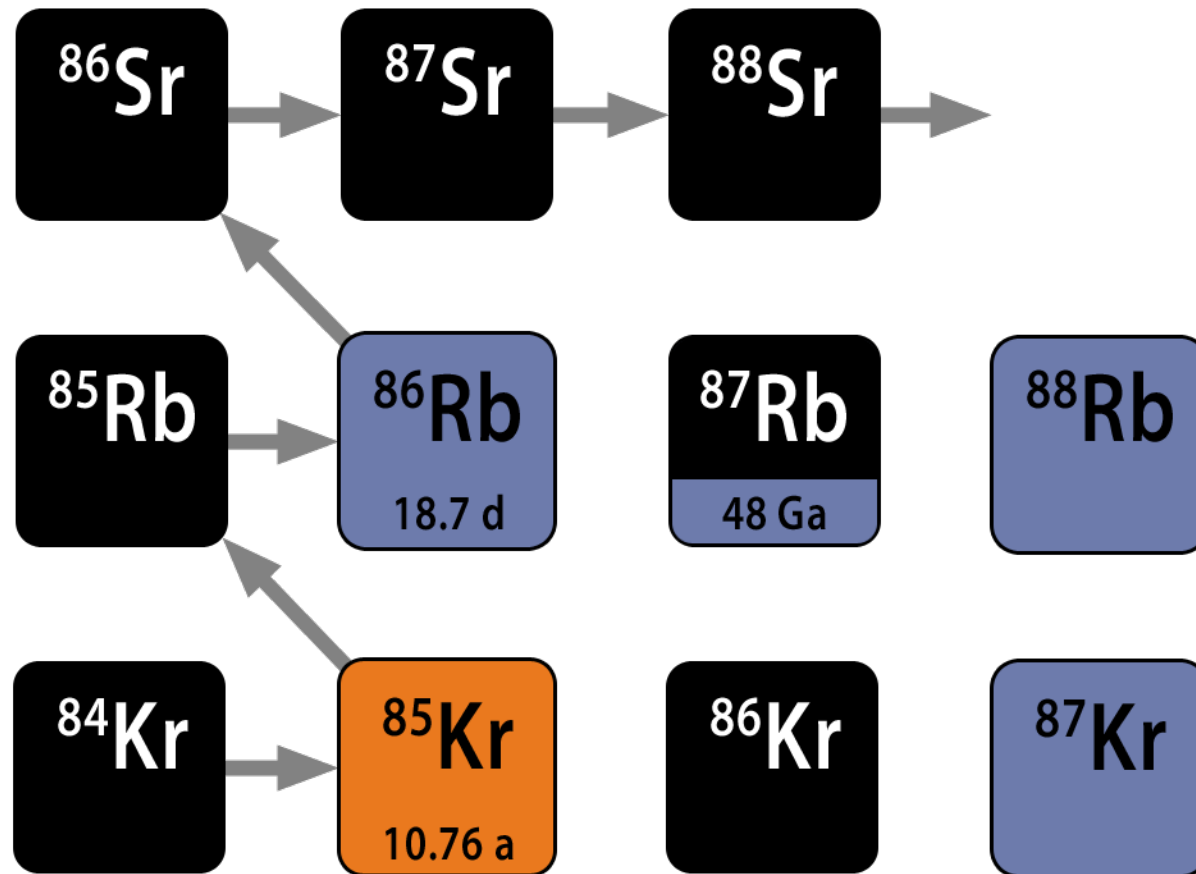
Meteorites and presolar grains



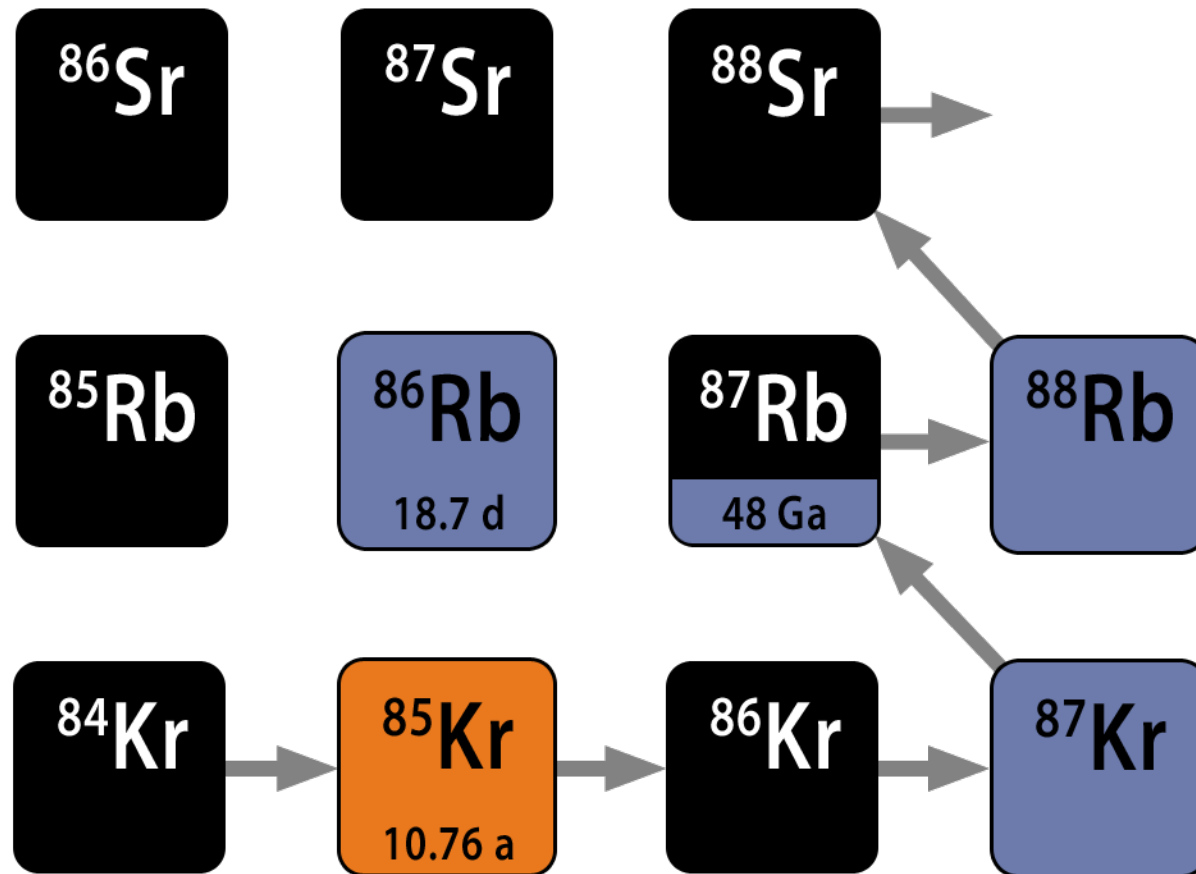
Meteorites and presolar grains



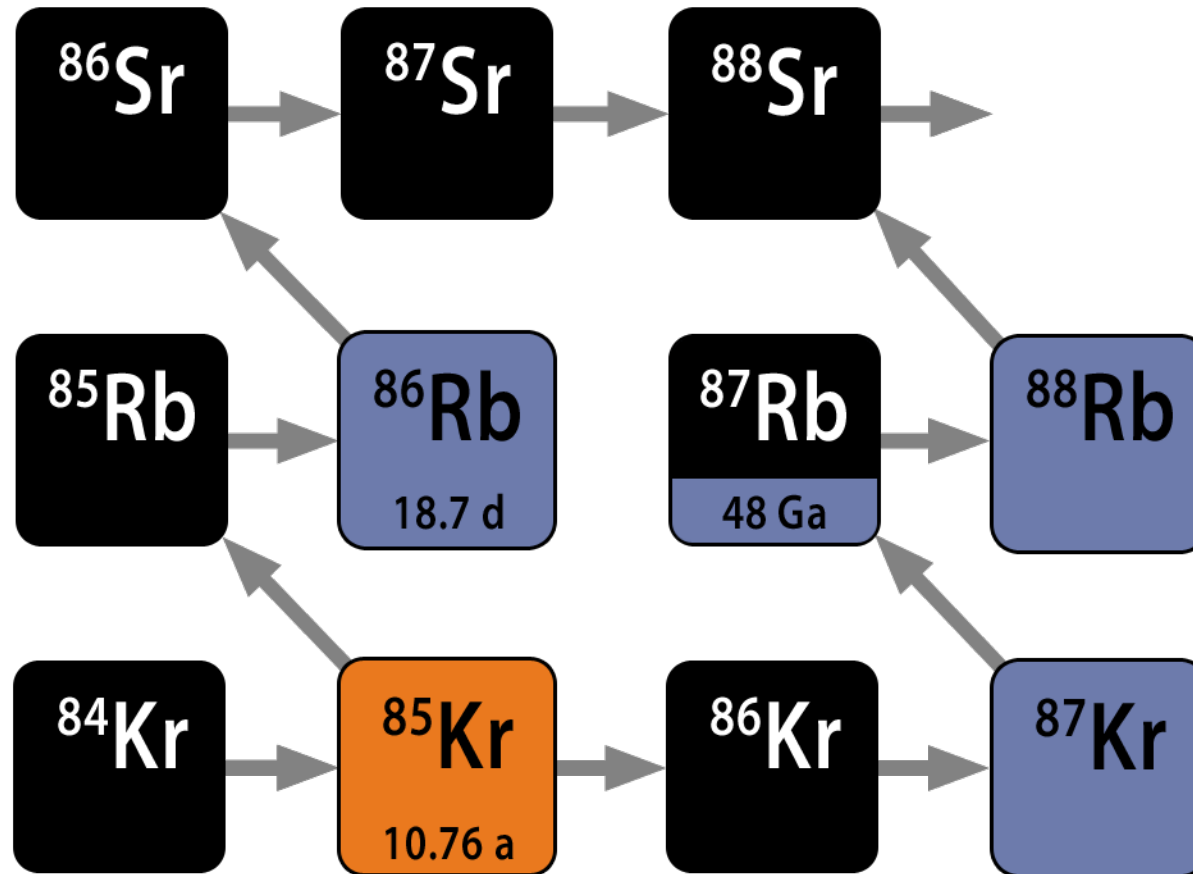
Branches in the synthesis paths



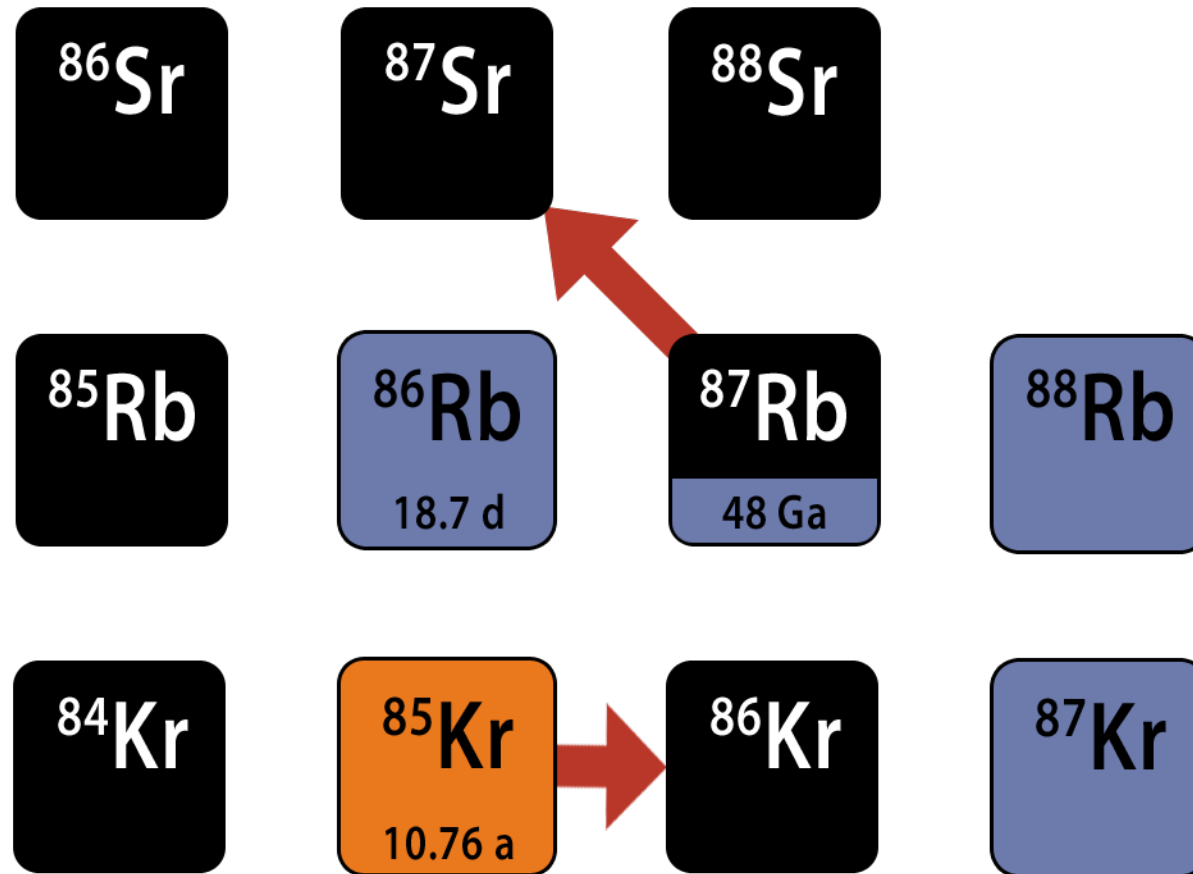
Branches in the synthesis paths



Branches in the synthesis paths

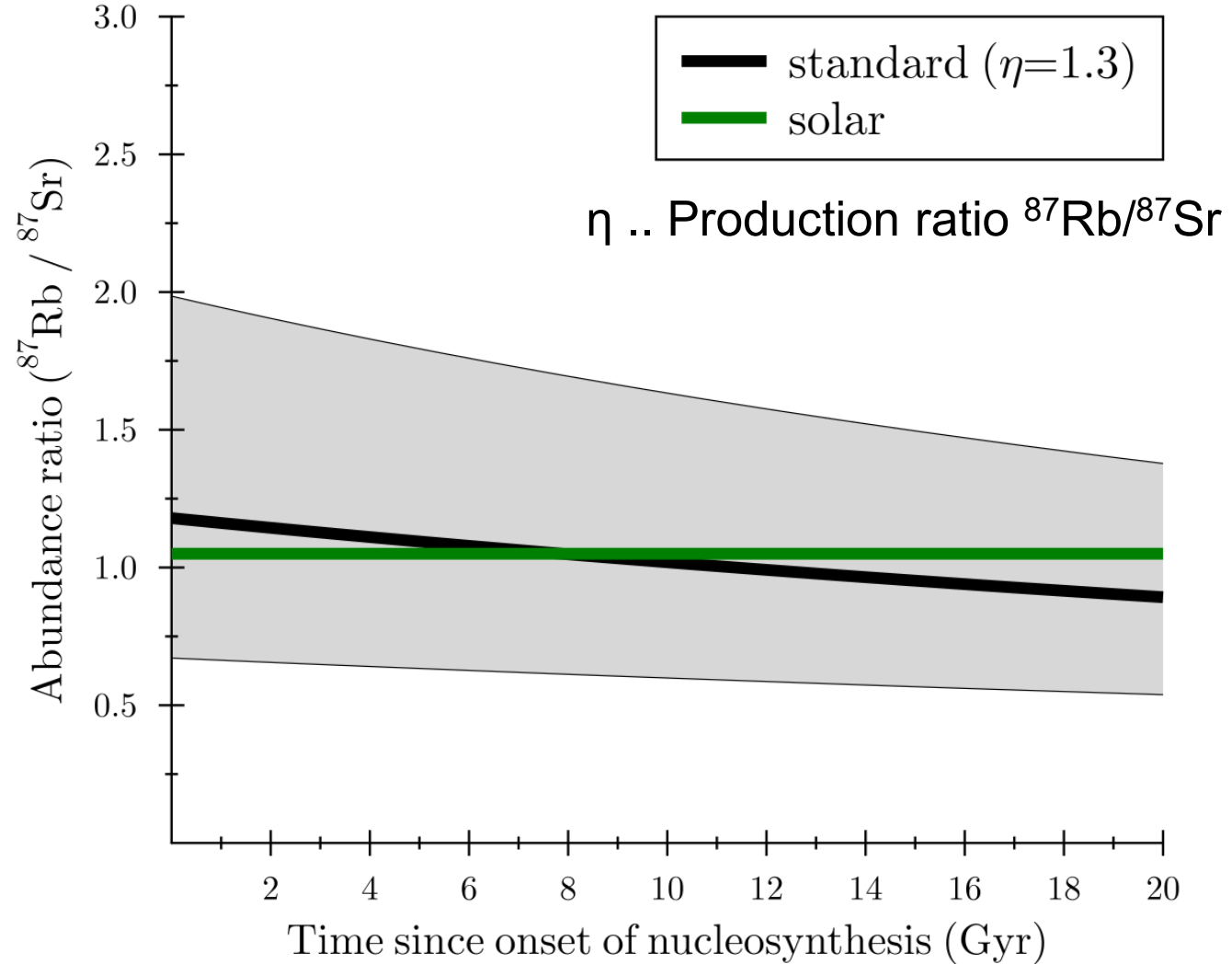


Branches in the synthesis paths



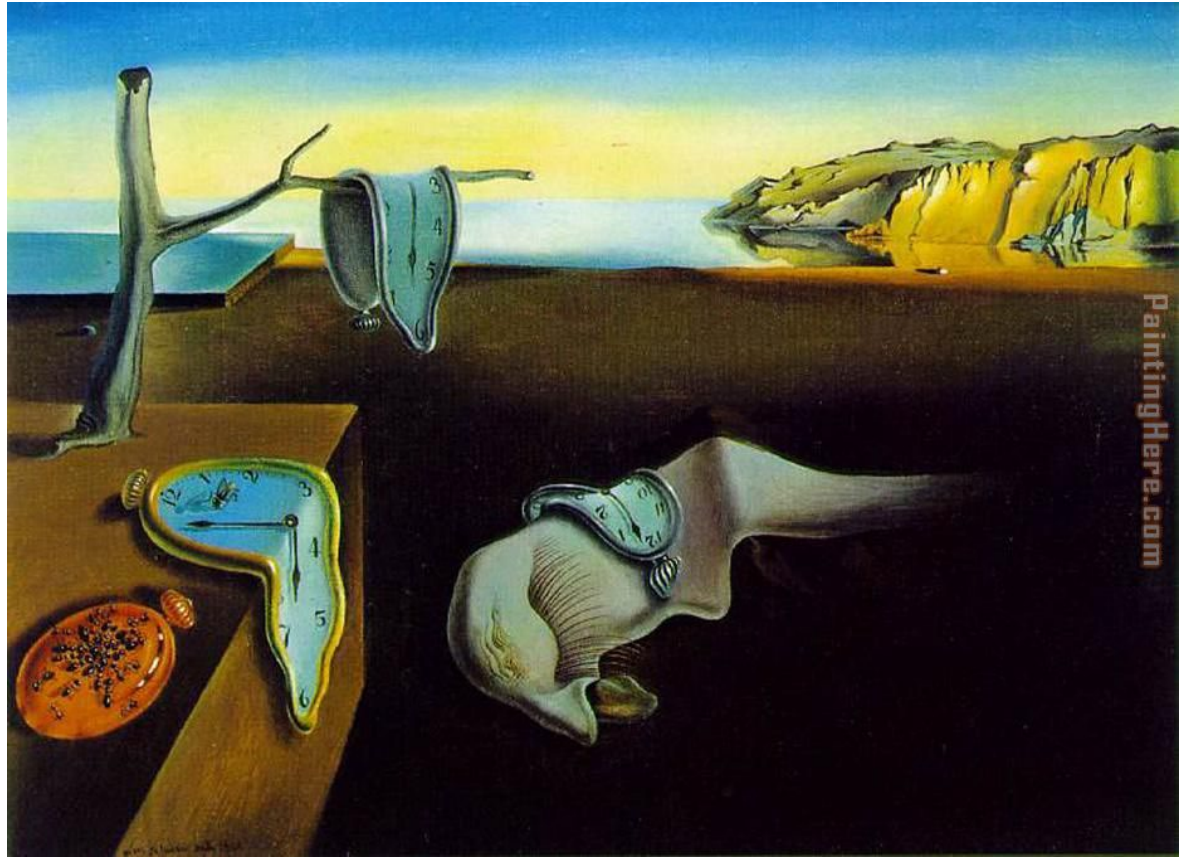
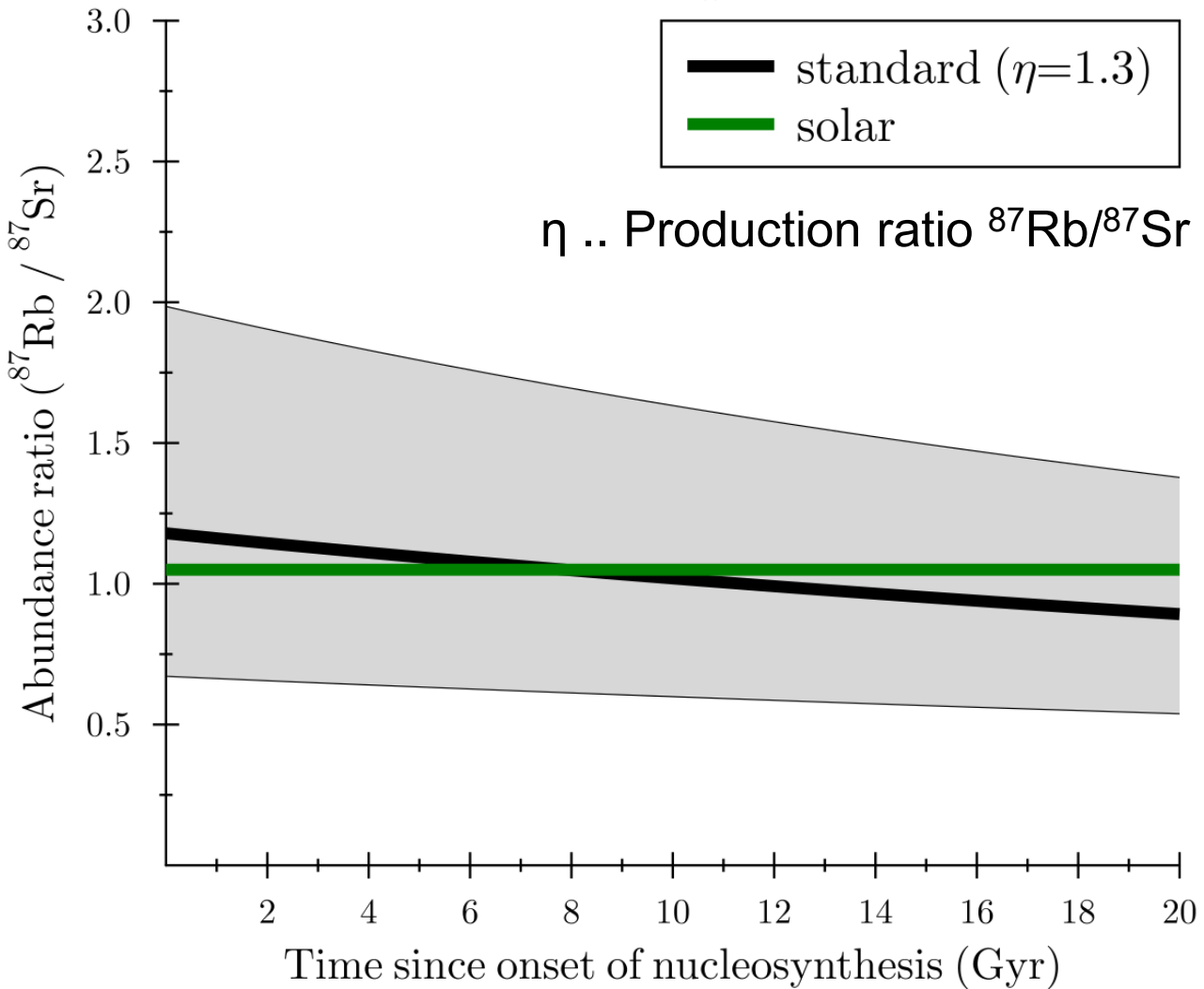
Cosmochronometer

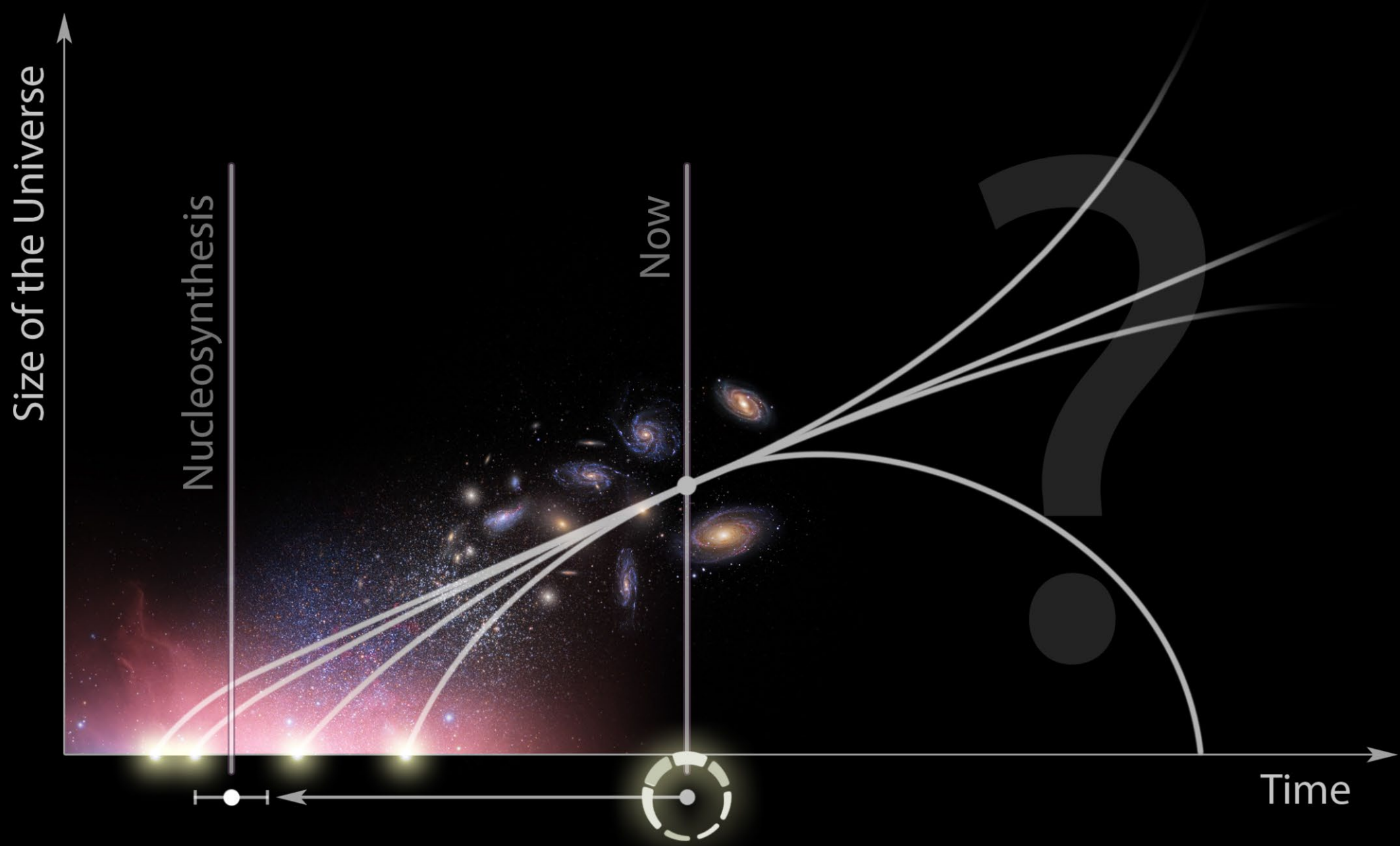
^{87}Rb and ^{87}Sr production



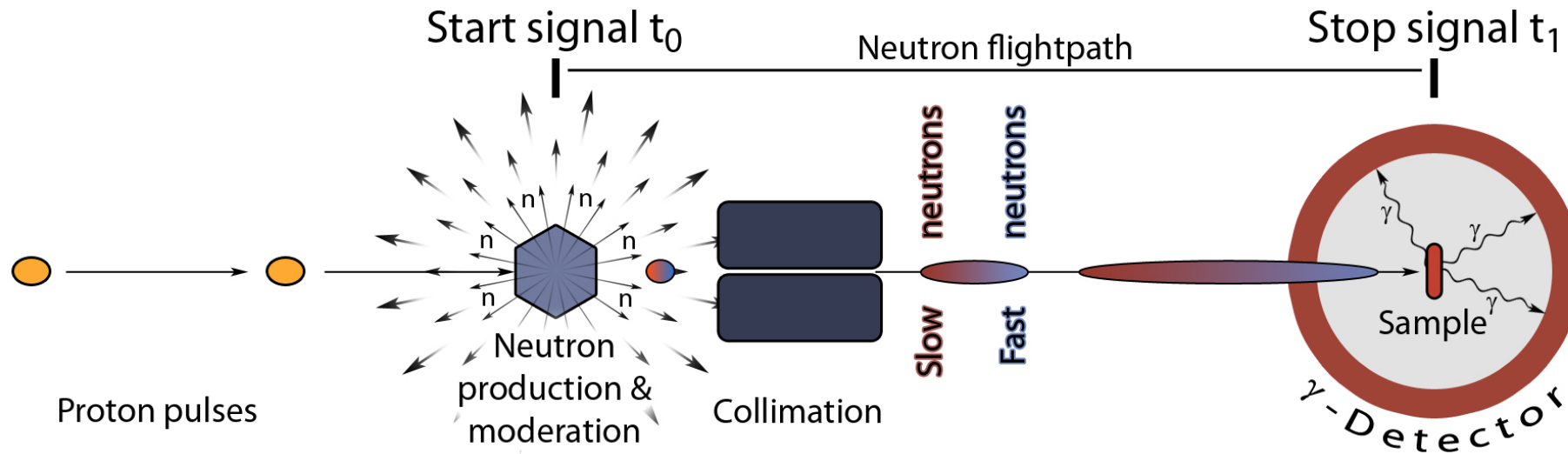
Cosmochronometer

^{87}Rb and ^{87}Sr production

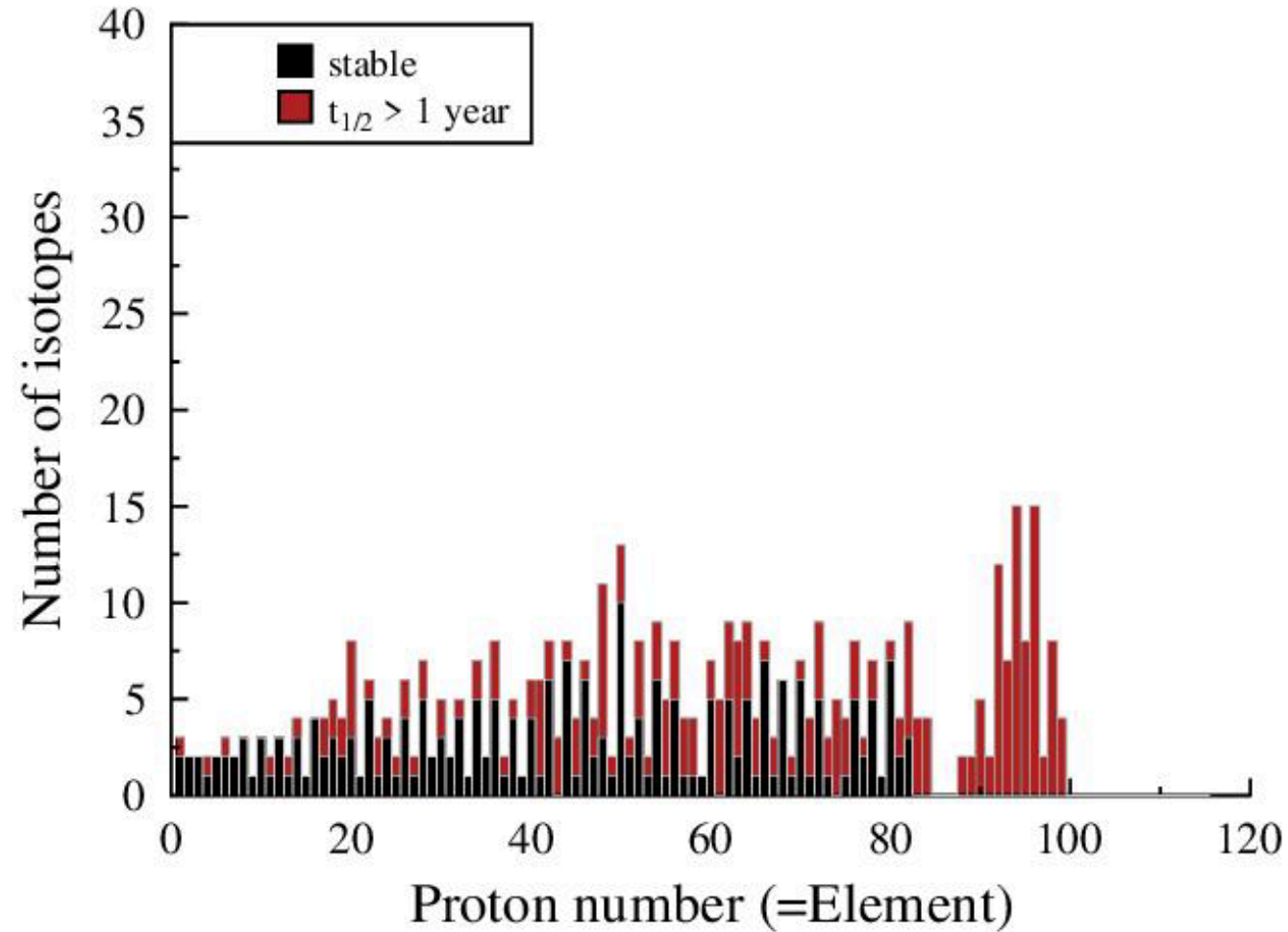




Neutron Reactions via time-of-flight technique



State of the art 2020 (e.g. **DANCE**)



Spallation-based neutron target

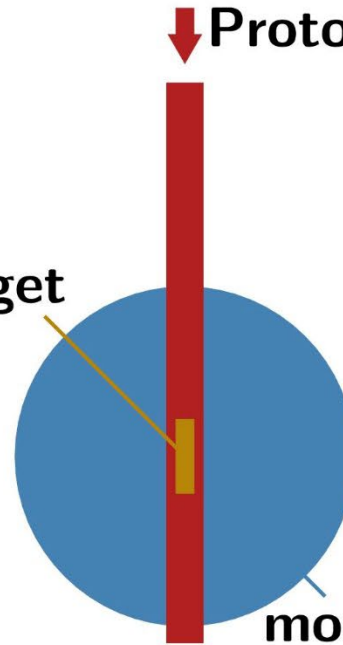
Tungsten spallation target

↓ Protons



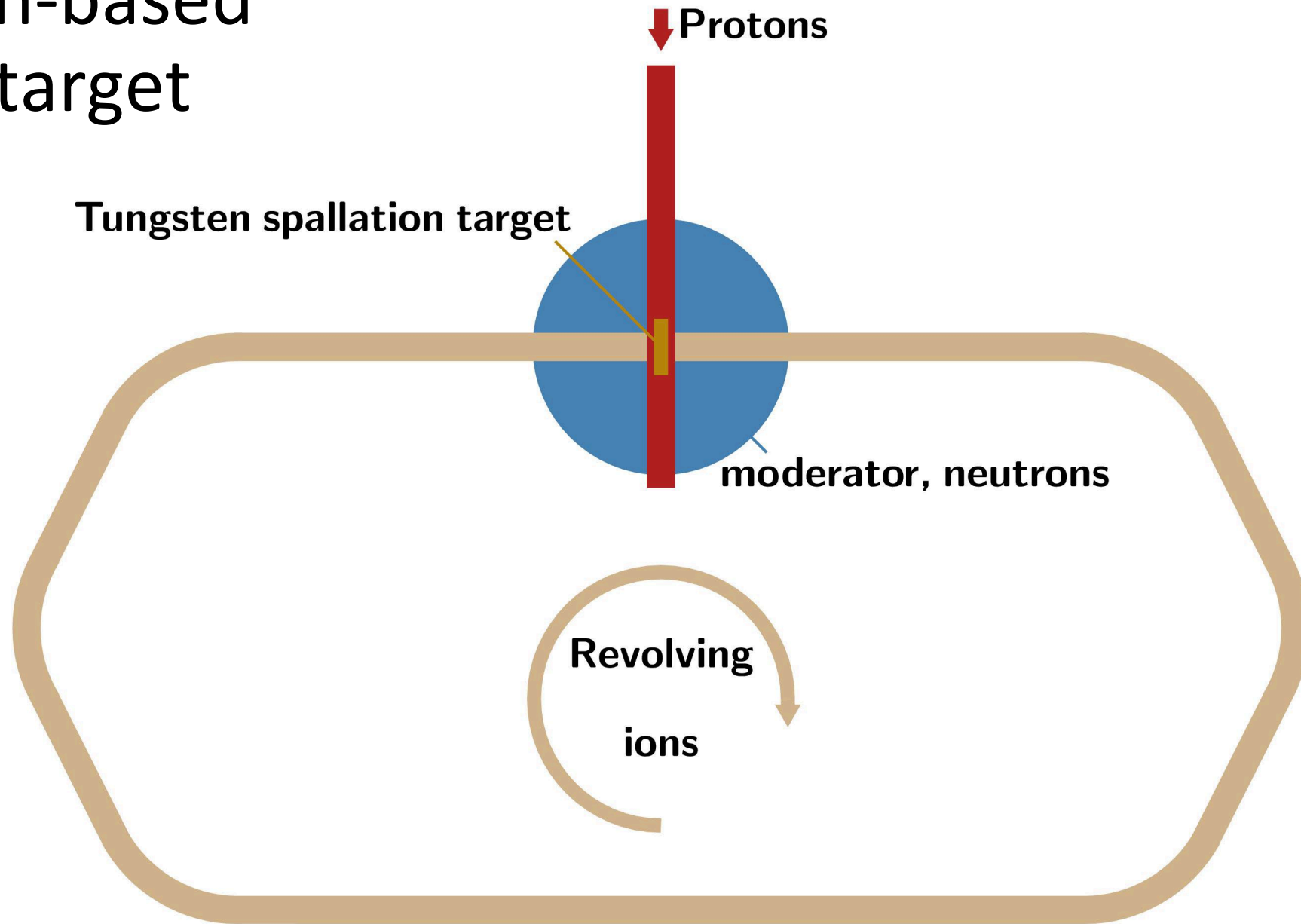
Spallation-based neutron target

Tungsten spallation target

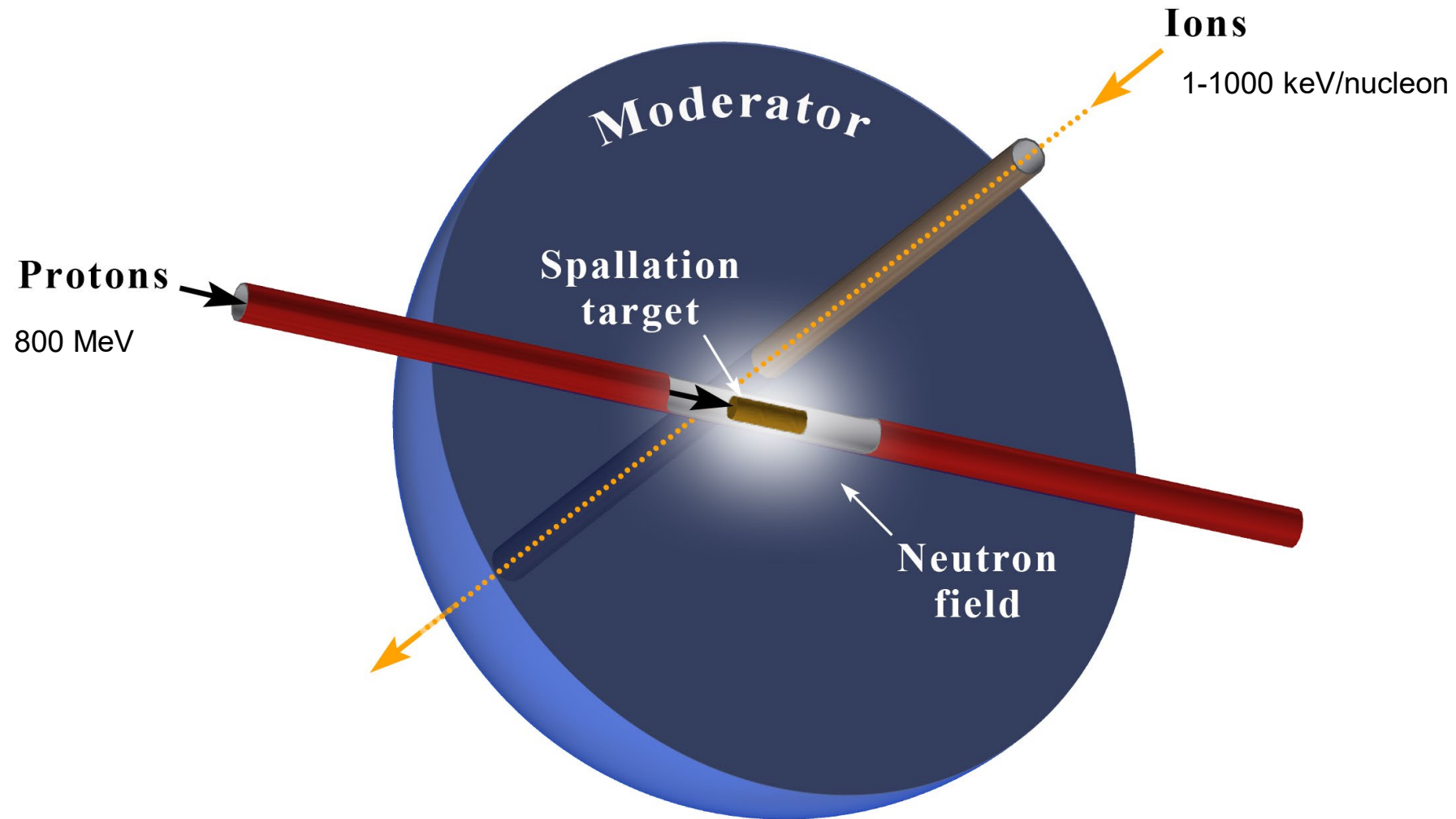


moderator, neutrons

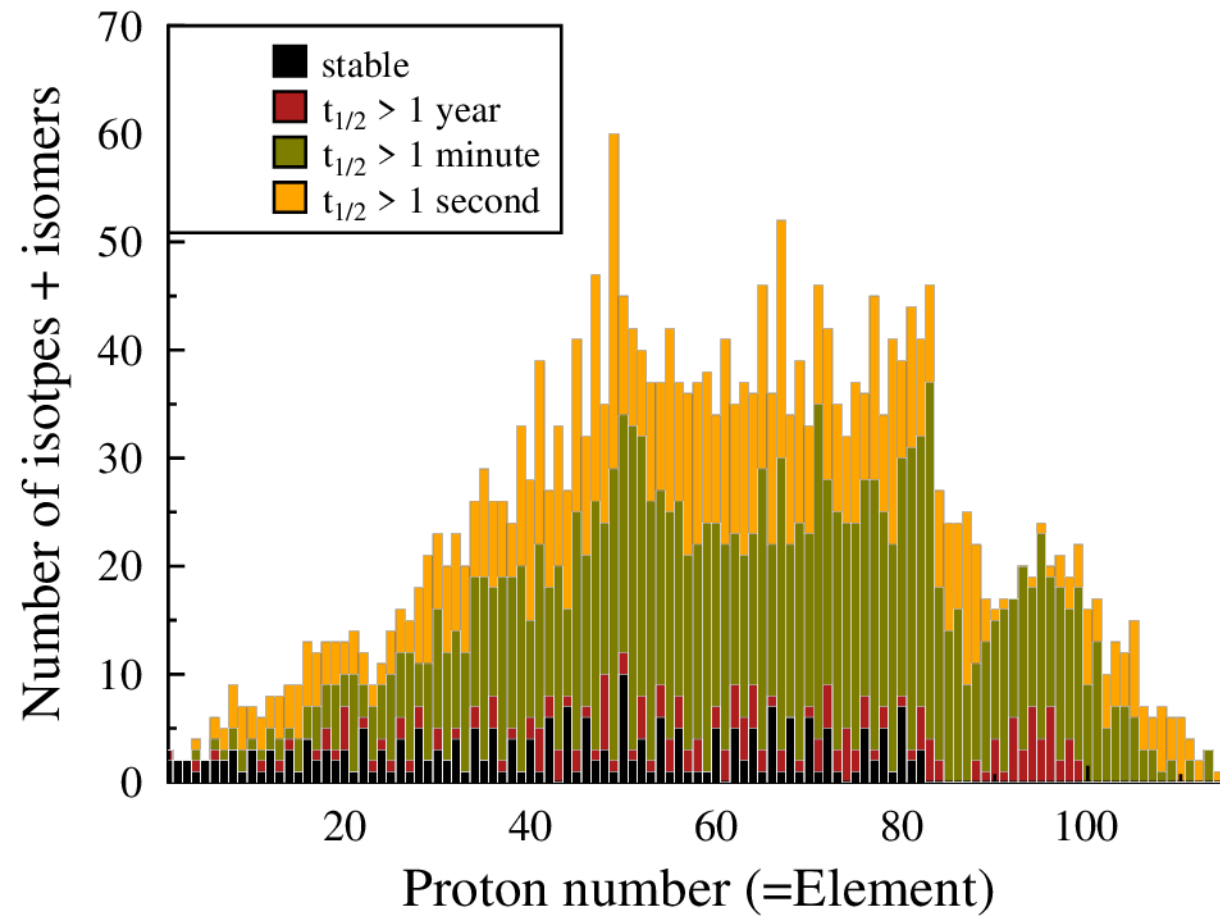
Spallation-based neutron target



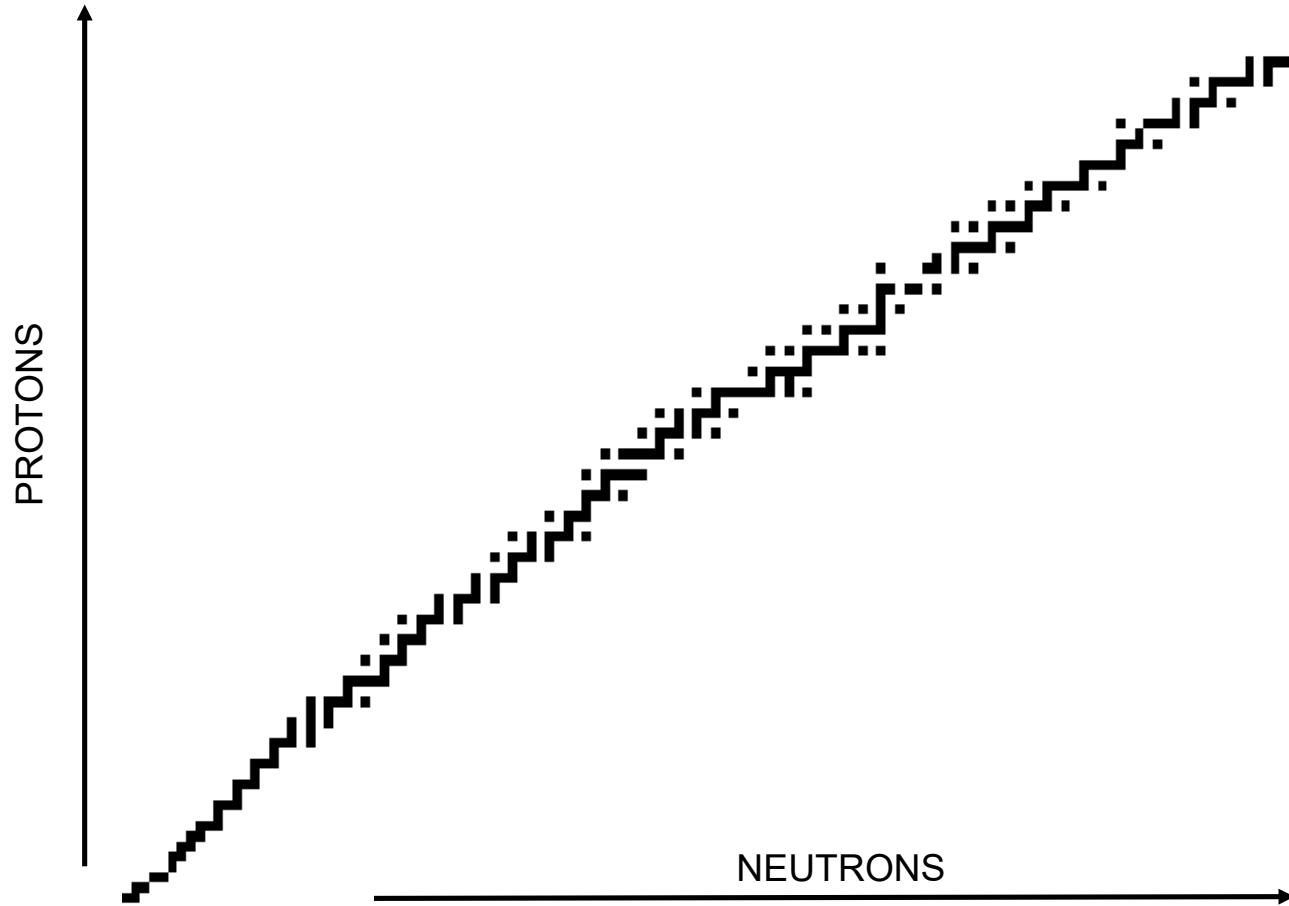
Proposed setup



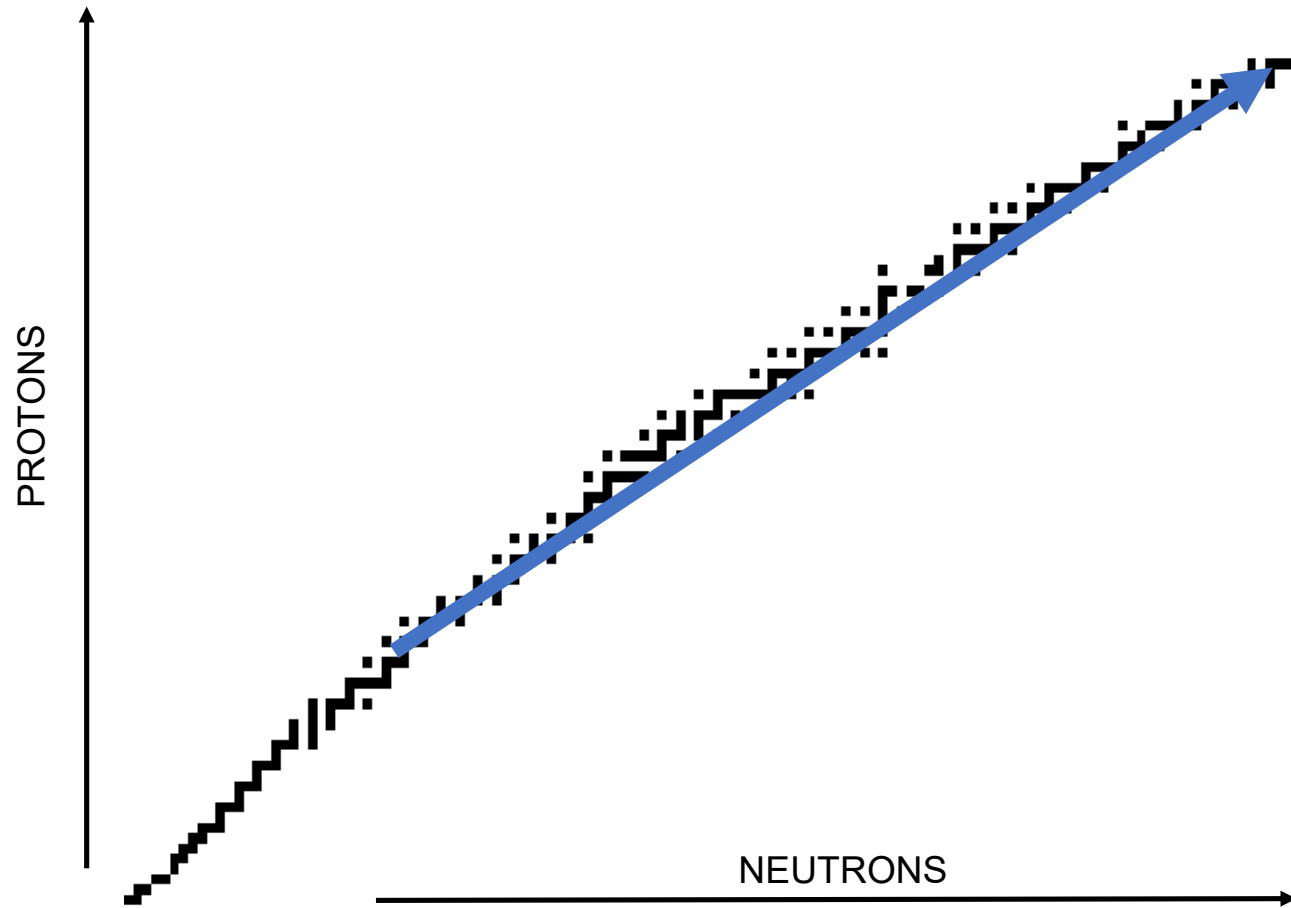
State of the art 2035? (N-TARGET+RING+ISOL)



The production of stable nuclei

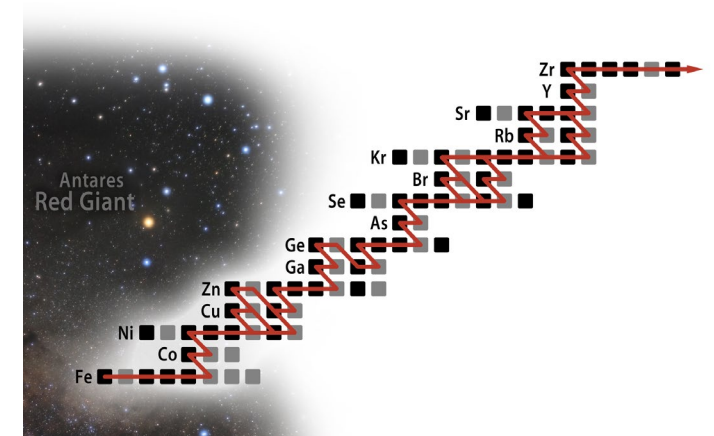


The production of stable nuclei

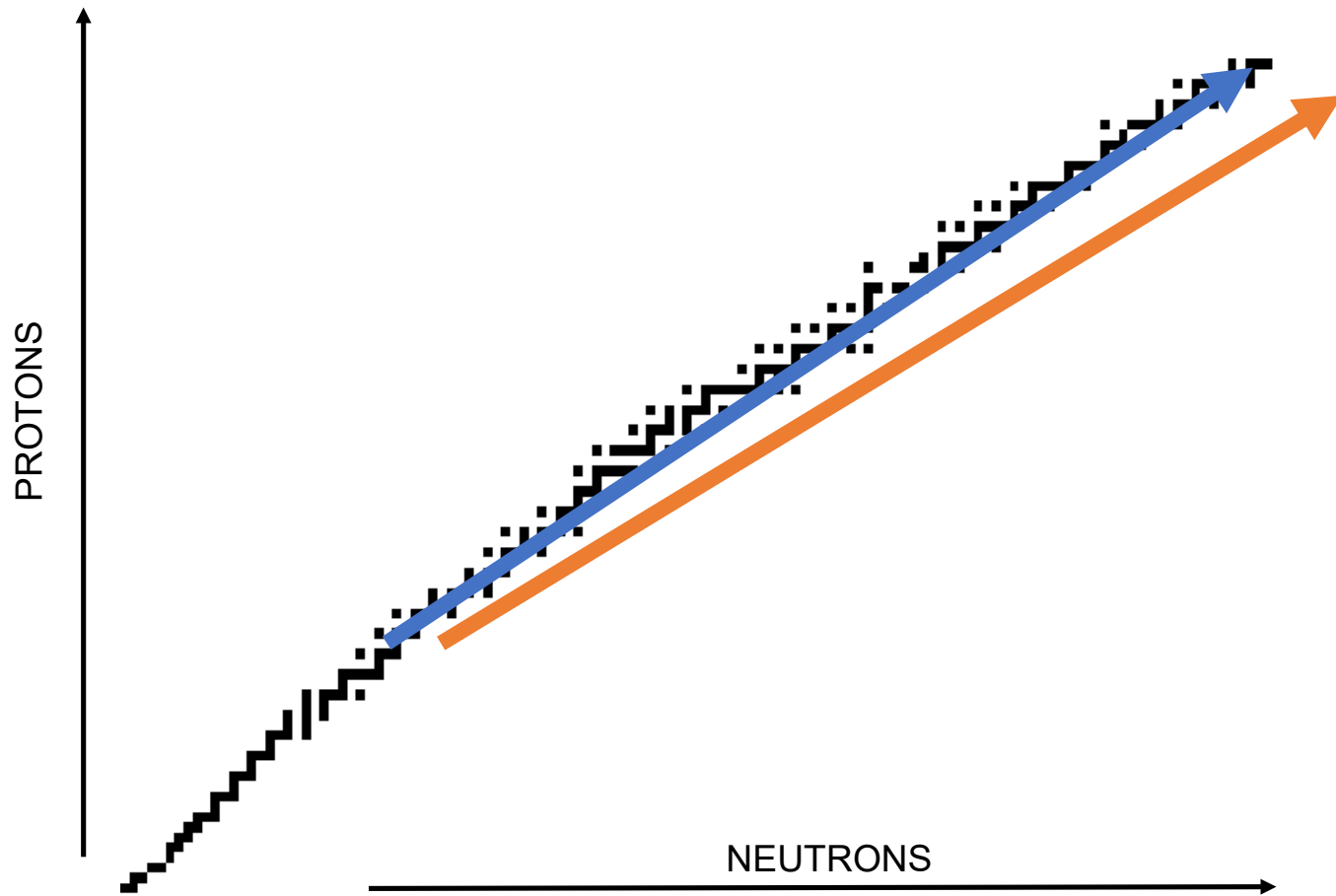


s-process:

completely covered



The production of stable nuclei

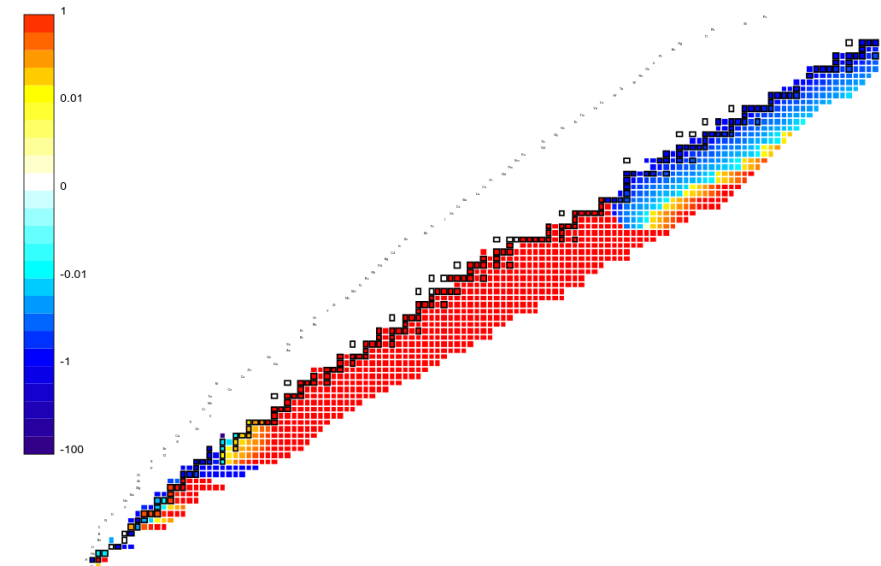


s-process:

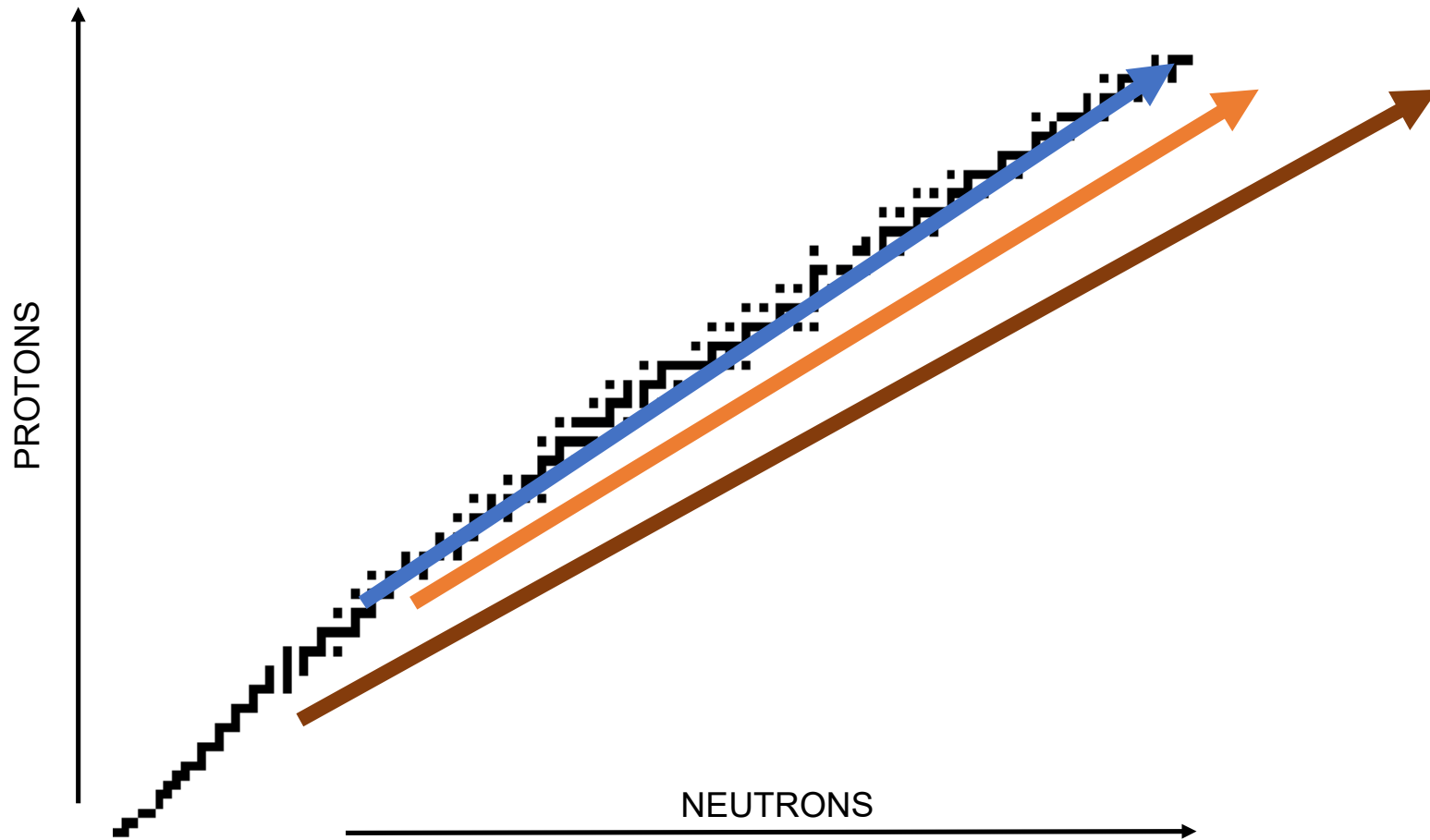
i-process:

completely covered

bottlenecks covered



The production of stable nuclei



s-process:

i-process:

r-process:

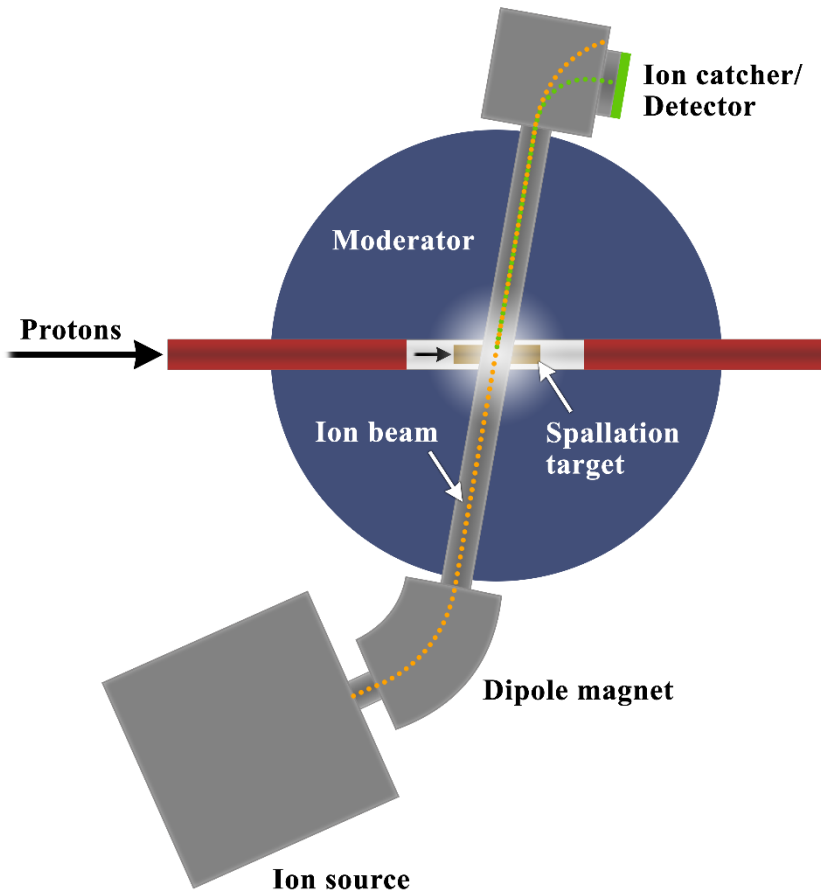
completely covered

bottlenecks covered

freeze out covered

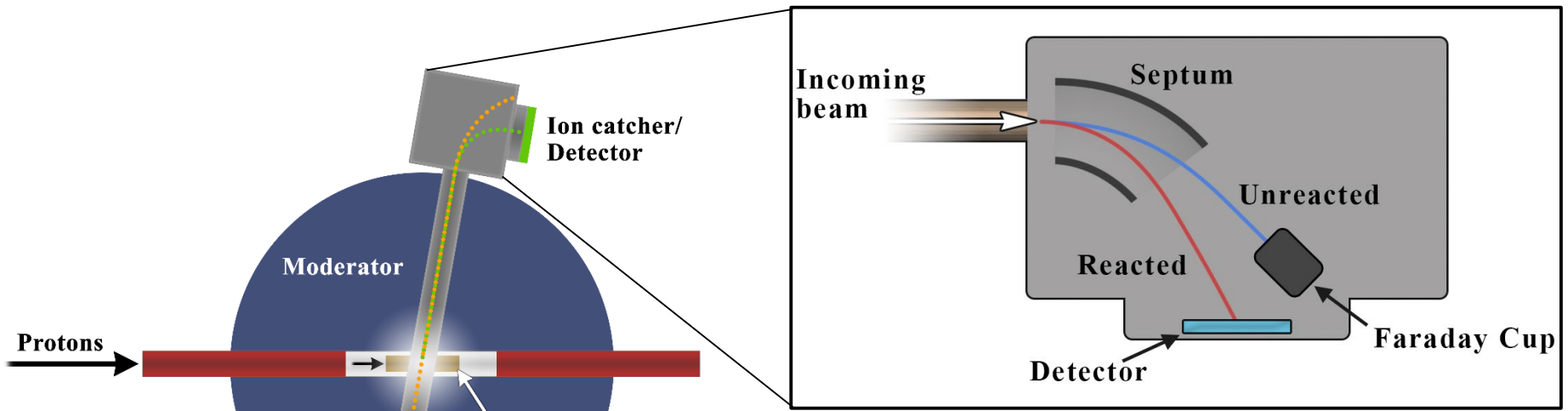


First stage: A neutron target demonstrator (LDRD-DR)

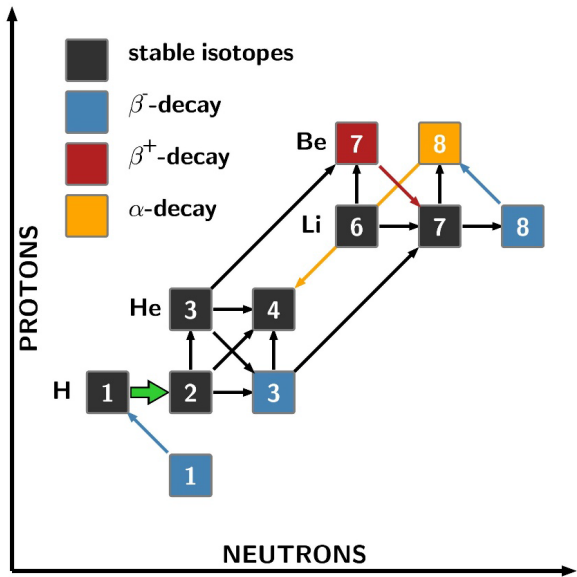


- Single pass experiment
- Blue room
- ^{84}Kr beam
- Detect decay of ^{85}Kr

Second stage: $n+p \rightarrow D$



- Single pass experiment
- Area A (higher n-density)
- ^1H beam
- Detect ^2H



Summary

